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ASIAN DEVELOPMENT BANK

A GUIDEBOOK ON MAPPING POVERTY THROUGH DATA INTEGRATION AND ARTIFICIAL INTELLIGENCE

APRIL 2021

ASIAN DEVELOPMENT BANK

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Note: In this publication, "\$" refers to United States dollars.

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FOREWORD

Since the Sustainable Development Goals (SDGs) were launched in 2015, both traditional and innovative
Stypes of data have become imperative in understanding the progress that has been made in achieving those goals. By providing more timely, granular, and comprehensive information, innovative sources complement traditional ones that are often constrained by high data collection costs. Conventional household or enterprise surveys, for instance, constitute a major data source for SDGs, but these often have sample sizes too small to provide enough granularity for highly targeted analyses. High costs also mean that these surveys are conducted too infrequently for timely measurement of indicators. On the other hand, conventional surveys and censuses serve as quality benchmarks for representativeness of data and adherence to statistical principles and standards that enable reliable inferences.

Indeed, to obtain timely, granular, and credible data entails integrating traditional with innovative data sources. Poverty statistics is an area where there have been several initiatives to blend multiple types of data. One noteworthy initiative involves using satellite imagery to provide more geographically disaggregated data than those published by government agencies. This approach leverages state-of-the-art computer imaging techniques to predict specific development indicators based on features on the ground.

The Asian Development Bank (ADB) designed a knowledge and support technical assistance called Data for Development in 2017 that aims to strengthen the capacity of national statistics offices to meet the increasing data demands for policymaking and monitoring of development goals and targets. One of its components focuses on subnational disaggregation of SDG indicators, particularly poverty statistics, that draws from recent studies combining geospatial data, satellite imagery, and powerful machine learning algorithms with traditional data sources and conventional methods to estimate the magnitude of poverty in specific locations. Such data are critical in aiding government and development agencies to distribute social assistance more efficiently. In the study, statisticians from ADB's Statistics and Data Innovation Unit within the Economic Research and Regional Cooperation Department worked with the Philippine Statistics Authority, National Statistical Office of Thailand, and World Data Lab to examine the feasibility of poverty mapping using satellite imagery and geospatial data.

This guidebook documents the study's key approaches step-by-step. It serves as a valuable reference for national statistics offices on how to use easily accessible resources such as satellite imagery to enhance the compilation of poverty statistics. The Key Indicators for Asia and the Pacific Special Supplement 2020 is recommended reading for users of this guidebook. The publication team was led by Arturo Martinez Jr. and Ron Lester Durante, under the overall direction of Elaine Tan. It was written by Ron Lester Durante, Arturo Martinez Jr., Mildred Addawe, Marymell Martillan, Joseph Bulan, Tomas Sako, and Martin Hofer, with valuable research and technical support from Katharina Fenz and Thomas Mitterling. Iva Lohovska from World Data Lab also provided insightful feedback on improving the guidebook, while Ma. Roselia Babalo, Rose Anne Dumayas, Raymond Adofina, and Ephraim Cuya provided operational support through its preparation. The cover of this publication was designed by Francis Manio. Manuscript editing

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was performed by Raynal Squires, while the publication's layout, page design, and typesetting were carried out by Judy Yñiguez.

We hope this guidebook will serve as a useful reference for national statistics offices across Asia and the Pacific in mapping the spatial distribution of poverty using a combination of traditional and innovative data sources.

 $\frac{1}{\sqrt{2}}$

Yasuyuki Sawada Chief Economist and Director General Economic Research and Regional Cooperation Department Asian Development Bank

ABBREVIATIONS

INTRODUCTION 1

Properly compiled data in poverty statistics provides visibility for socioeconomically disadvantaged people in society. It sheds light on their demographic profiles, their magnitude, location, and their needs, all of which are critical inputs for the design of interventions in a development agenda.

In developing countries, poverty statistics are typically derived from household surveys designed to provide reliable estimates at national, regional, provincial, or other highly aggregated levels. However, as better disaggregated data can facilitate more effective targeting of socioeconomic programs, it is important to explore alternative data sources that can complement these surveys.

Satellite imagery is a potentially useful source of alternative data which may be used to enhance the granularity of poverty statistics compiled from household surveys. The emergence of satellite data has invigorated efforts to measure poverty on a gridded level from space. A novel approach entails using artificial intelligence to predict the prevalence of poverty (or other indicators) based on satellite image features.¹ Since data from images are naturally unstructured, noisy, and difficult to process statistically, one can design computer vision techniques to extract patterns that may be used to associate them with poverty.

Mapping Poverty through Data Integration and Artificial Intelligence: A Special Supplement of the Key Indicators for Asia and the Pacific, a report published by the Asian Development Bank (ADB), documents the results of using computer vision techniques to map the spatial distribution of poverty in the Philippines and Thailand.² The country-specific reports, Mapping the Spatial Distribution of Poverty Using Satellite Imagery in the Philippines and in Thailand, provide more detailed discussion on the methodology.³ The first step of the methodology entails training a convolutional neural network (CNN)—an advanced type of machine learning algorithm commonly used for image classification-related tasks—to predict nighttime light data using daytime images as input. Intensity of lights at night is a good proxy for wealth and human interaction on the ground and this kind of abundant, granular information meets the high-volume data requirement for training machine learning algorithms. In the process of learning to "predict" nighttime light intensity, the CNN learns to detect general features in images, or latent variables, related to light intensity that can be used for other tasks, like estimating poverty measures. To maintain consistency with published official statistics, the condensed, image-based information can be averaged on a coarser level to align with the level of information available in government-published poverty estimates. To speed up learning and reduce the amount of data needed for the process, a CNN that has already been trained on some image databases is used to assign labels to larger databases of images.

¹ N. Jean et al. 2016. Combining Satellite Imagery and Machine Learning to Predict Poverty. *Science*. 353 (6301). pp 790–794.

² Asian Development Bank (ADB). 2020. *Mapping Poverty through Data Integration and Artificial Intelligence: A Special Supplement of the Key Indicators for Asia and the Pacific*. Manila.

³ ADB. 2021. Mapping the Spatial Distribution of Poverty Using Satellite Imagery in the Philippines. Manila; and ADB. Forthcoming. *Mapping the Spatial Distribution of Poverty Using Satellite Imagery in Thailand*. Manila.

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In the second step, prediction of nighttime light intensity is discarded and the trained CNN alone is used to summarize the complex multidimensional input of image data into a single vector. This vector has hundreds of features, each assigned a single value in every image. These features are a representation of what the network detects in an image. They have several advantages over raw pixel values, most notably that convolutional layers scan over the image using kernels so that it does not matter where features are placed on the image.

To combine grid-based image features with survey-based poverty data, the value of each feature within the given survey areas is averaged. The final training step uses a ridge regression to find the relationship between the image features and survey-based poverty statistics. The trained CNN and ridge parameters can then be used to predict poverty using only a daytime image as input. The process is illustrated in Figure 1.

Figure 1: Road Map of Methodology for Predicting Poverty Using Satellite Imagery

Notes: The procedure requires three types of data: geographically disaggregated poverty statistics, daytime satellite imagery, and images of earth at night. After pre-processing and cleaning these data (Step 1), Step 2 trains an algorithm to classify (daytime) satellite images into different classes of night light intensity. Step 3 extracts the image features from the last layer of the trained algorithm. In Step 4, the image features are averaged so the space enclosed in grids corresponds to the level at which poverty-labeled images are available. These are regressed using the target variable of the survey to find the relationship between features and the target variable. Step 5 summarizes the full pipeline from input image to target variable.

Source: Graphics generated by the study team.

This guidebook outlines the step-by-step procedure summarized in Figure 1. The guidebook is intended as a one-stop reference for researchers and other development practitioners (particularly from national statistics offices) who wish to apply these methods for exploratory studies using tools that are readily accessible and without significant cost. Because we strongly believe in the straightforward methods and tools described here, other (sometimes proprietary) tools that may be more effective in conducting largerscale poverty mapping initiatives are not discussed.

Users of this guidebook are encouraged to first read the ADB report (footnote 2), particularly the section describing the methodology, before going through the step-by-step procedure outlined here. Users are also advised to check for updates to the software and services referred to and pictured in screenshots in this guidebook. The discussions in this guidebook are meant for educational purposes. It should be noted that trademarks of tools and resources used are owned solely by the respective developers, and this guidebook is not endorsed by or affiliated with these companies in any way.

HARDWARE AND SOFTWARE 2 HARDWARE AND SOFTWARI
2 REQUIREMENTS AND SETUP

Hardware

 Minimum system requirements: 1.6 gigahertz 4-core processor or better, 8 gigabytes (GB) RAM, 10 GB of free hard disk space with reliable internet connection

Software

- R version 4 or higher
- RStudio version 1.4 or higher
- R Packages: caret, fasterize, gdalUtilities, mclust, raster, rasterVis, sf, tidyverse
- Google Chrome browser version 79.0.3945 or higher
- Fastai Python library version 1.0.61
- Gmail account, Google Drive with at least 5 GB free space, and Google Earth Engine account

Software Requirement Setup

R and RStudio

For step-by-step procedure in downloading and installing R and Rstudio, refer to this page: [https://rstudio-education.github.io/hopr/starting.html.](https://rstudio-education.github.io/hopr/starting.html)

Installing Rtools

Rtools is used to build R and R packages because some of the packages are downloaded as source code and need to be compiled.

For information on how to install and test Rtools, refer to this page: [https://cran.r-project.org/bin/windows/Rtools/.](https://cran.r-project.org/bin/windows/Rtools/)

Installing R packages

The required packages are caret, fasterize, gdalUtilities, mclust, raster, rasterVis, sf, and tidyverse. Table 1 provides a description of these packages.

To install these packages, type the following commands in the Source Panel:

```
install.packages(c("caret",
                    "fasterize",
                    "gdalUtilities", 
                    "mclust",
                   "raster",
                   "rasterVis", 
                   "sf",
                   "tidyverse"), 
                dependencies = T)
```
Then click the $\|\cdot\|$ Source $\cdot\|$ icon to execute the entire script.

Table 1: Description of Required R Packages

Some of the packages and/or their dependencies need to be installed from source through the help of Rtools. A dialog box will ask permission to install packages from source.

Click **Yes** to start package download and installation.

The Console Panel will revert to prompt once all packages are installed. Review the **Console Panel** outputs to check for errors in package installations.

Chrome Browser

Install Google Chrome Web Browser version 79.0.3945 or higher.

For step-by-step procedure in downloading and installing Google Chrome, refer to this page: <https://support.google.com/chrome/answer/95346>.

Google Account

Setting up a new Google account.

For step-by-step procedure in creating a Google account refer to this page: [https://support.google.com/](https://support.google.com/accounts/answer/27441%3Fhl%3Den%23) [accounts/answer/27441?hl=en#](https://support.google.com/accounts/answer/27441%3Fhl%3Den%23).

If you prefer to use an already existing Google account, verify that its associated Google Drive has at least 5 GB of free storage space.

Google Earth Engine

Google Earth Engine (GEE) is a cloud-based geospatial processing tool with built-in spatial datasets that goes back more than 4 decades. A sign-up is required using an active Google account to use the GEE service.

Refer to this page to sign up and get access for Google Earth Engine: [https://signup.earthengine.google.com/.](https://signup.earthengine.google.com/)

Below is the Google Earth Engine Code Editor.

Geometry tools

DATA PREPARATION 3

Daytime Satellite Imagery Processing

Downloading the Shapefiles

A shapefile is a simple vector data storage format for storing the location, shape, and attributes of geographic features. The geographic features in a shapefile can be represented by points, lines, or polygons.4 Shapefiles determine the extent of satellite imagery to download. The administrative boundaries of the shapefiles should be consistent with official statistical data.

Shapefiles can be downloaded from various sources, but the most common are the Humanitarian Data Exchange (HDX) (*www.humdata.org*) and Database of Global Administrative Areas (GADM) (*www.gadm.org*).

HDX is an open platform for sharing data across crises and organizations. Launched in July 2014 by the United Nations Office for the Coordination of Humanitarian Affairs, HDX aims to make humanitarian data easy to find and use for analysis. HDX shapefiles are derived from original datasets sourced from relevant government agencies (e.g., national statistics offices, mapping agencies) and attached with standard geographic codes. These shapefiles have been vetted, configured, and provided with live services by the Information Technology Outreach Services of the Carl Vinson Institute of Government - University of Georgia. These shapefiles are also updated every year.

GADM is a high-resolution database of country administrative areas that provides maps and spatial data for all countries and their subdivisions. The current version is 3.6, which delimits 386,735 administrative

⁴ Environmental Systems Research Institute (ESRI). 1998. ESRI Shapefile Technical Description: An ESRI White Paper – July 1998.<https://www.esri.com/Library/Whitepapers/Pdfs/Shapefile.pdf>.

areas with high spatial resolution and an extensive set of attributes. One limitation of using GADM is that the administrative subdivisions could possibly differ on a country basis.

For the following steps, Thailand files are used for illustration.

STEP 1

In the browser address bar, type the HDX web address, www.humdata.org, and press **Enter**. From the top bar, click **Search Datasets.** Type **<country_name> administrative boundary**. For this illustration, type **Thailand administrative boundary** and press **Enter**.

STEP 2

Click the link to the country's administrative boundary shapefile. Click **Thailand administrative levels 0-3 boundaries**.

Browse and select the country-level shapefile and the administrative boundary shapefile coinciding with the published poverty estimates.

For this illustration, select **tha_adm_rtsd_itos_20190221_SHP_PART_2.zip**. Then click **Download.**

The shapefile is compressed in a ZIP file and automatically saved in the default download folder.

Open the Downloads folder. Extract the shapefile from the ZIP file. Check the information note attached to the ZIP file as different countries may have different notations.

In the case of Thailand, the following notations are used:

- adm0 Country level
- \blacksquare adm1 Provincial level
- \Box adm2 District level
- adm3 Sub-district level (tambon)

Generating Centroids for Satellite Imagery

For this illustration, municipal boundary shapefiles are used to generate grids from raster pixels. Then centroids are obtained for each grid. Outputs are saved as comma-separated values (CSV) file.

Grid centroids will be used to determine the center of the daytime satellite imagery tile to be downloaded. Each tile will serve as input image for training the CNN model.

STEP 1

Open RStudio.

STEP 2

Click the **Open File** icon **in the toolbar.**

Search the R code: *grid_cell_selection.R* and click **Open**.

The administrative boundary shapefiles that correspond to the geographical level of the published poverty data will be used to generate grids from raster pixels. Obtain the centroids of each grid. Then generate the output as CSV file.

STEP 3

Load the R packages by typing *library(package)*. On the R console window, type the following commands and press **Enter**.

- **sf** is for interpreting and operations on vector shapefiles
- **raster** is for raster object operations
- **fasterize** is for rasterizing vectors
- **tidyverse** is for data manipulation

Select the working directory (i.e., the active computer folder) using the function **tk_choose.dir()** from the package **tcltk** (tcltk is a built-in package that provides the GUI for R; this command opens a window for selecting the target folder).

```
9 \cdot # set working directory----
10 wd \leftarrow \text{tdt}: tk_choose.dir(caption = "Select Working Directory")
11 setwd(wd)
```
Set the working directory by typing *setwd()*.

STEP 5

Set the code pertaining to the country of study by typing *country = "code"*.

STEP 6

Calculate the grid size.

Grid size is the product of the satellite resolution (i.e., satellite granularity in meters/pixel) and the CNN input image size (i.e., set.grid.resolution.px in pixels).

Most of the CNN architecture is trained on ImageNet ([http://www.image-net.org/\)](http://www.image-net.org/), which is a database of human labeled images, like ResNet, which uses 256x256 pixel resolution. Though most have image input size of 224x224 pixels, these architectures can also benefit from higher resolution images such as 512x512 pixels, 1024x1024 pixels or higher. However, this increase in resolution also increases the file size of each image, constraining the graphics processing unit's (GPU) memory where it will be stored and processed during the CNN training process. The higher the resolution, the longer the training period since you may need to train the model in smaller batches of images.

Satellite granularity was based on Landsat's⁵ resolution of **15 meters/pixel** after pansharpening.

Landsat is the longest running program for acquisition of satellite imagery of Earth.

```
# in pixel, image size required by CNN model
30
    set.grid.resolution.px <-256satelite.granularity \langle -15 \rangle # in meter/pixel,
31
32
                                      # Landsat resolution after pansharpening
33
34 gridsize \leftarrow set.grid.resolution.px * satelite.granularity
```
The grid size is equal to 3840 meters.

Landsat is used as reference for grid computation because it has lower resolution (i.e., larger pixel size), hence, more coverage and image detail. For the higher resolution Sentinel 2 satellite, more pixels can be derived for the same grid size.

STEP 7

Select the file path of the administrative boundary shapefile that is consistent with the granularity of the government-published estimates. Use the function **tk_choose.files()** to refer to GUI-based file selection.

```
36 * # Select location of administrative boundary shapefile----
37 shp_path <- tcltk::tk_choose.files(filters = matrix(c("SHP",".shp","All files","*"),2,2,byrow = T),
38
                                        caption = "Select Admin Boundary Shapefile")
39
40 + # read shapefile to see which points are within the country border----
41 ADM_sf <- read_sf(shp_path)
```
Next, load the shapefile using *sf* function's *read_sf()*.

STEP 8

Create a new column containing the numeric portion of the administrative boundaries' geographic code. The shapefile's PCODE usually contains a country code prefix. Thus, use a *stringr* package's **str_extact()** function to get only the numeric portion of ADM3_PCODE entries.

```
45 # we use stringr package's str_extact() function to get the numeric portion of ADM3_PCODE entries
46 ADM_sf$geocode <- as.numeric(str_extract(ADM_sf$ADM3_PCODE,"[0-9]+\\.*[0-9]*"))
```


The Coordinate Reference System (CRS) is a system used to define the position on the earth's surface. It allows merging of spatial datasets accurately and facilitates calculation of distance and surface area properly. There are two types of CRS: the Geographic Coordinate System (GCS) and the Projected Coordinate System (PCS). GCS covers the entire globe, while PCS is localized to lessen visual distortion in a specific region. GCS is based on sphere coordinates and utilizes angular units (e.g., degrees, minutes, seconds), while PCS is plane-based and uses linear units (e.g., meter, feet). World Geodetic System 1984 (WGS84) is an example of GCS. Universal Transverse Mercator (UTM) is an example of PCS.

Define the CRS variables in Proj.4 format. There are several websites that host Proj.4 CRS of different projections, two of which are <https://spatialreference.org/> and <https://epsg.io/>. Use the CRS to transform the shapefiles from GCS into PCS. Make sure to check the appropriate PCS for the country of study.

Type the following commands and press **Enter**.

```
49 - # Define crs variables ----
50 # There are several websites that hosts Proj.4 CRS of different projections,
51 # two of which are https://spatialreference.org/ and https://epsg.io/
S2 WGS84 <- "+proj=longlat +datum=WGS84 +no_defs +ellps=WGS84 +towgs84=0,0,0"<br>53 UTM_CRS <- "+proj=utm +zone=47 +datum=WGS84 +units=m +no_defs" #Thailand is located at zone 47N
```
STEP 10

Check the projection information of the shapefile to verify its CRS.

56 # check the projection information of the shapefile 57 $print(crs(ADM_s f))$

> print(crs(ADM_sf)) CRS arguments: +proj=longlat +datum=WGS84 +no_defs

Transform the shapefile from GCS to PCS. Use **sf** package's **st_transform( )** to change the shapefile's CRS.

```
59
    # transform shapefile from WGS84 to UTM
    ADM_UTM_sf <- st_transform(ADM_sf,UTM_CRS)
60
61
62 # check the projection information of the shapefile to verify CRS
63 print(crs(ADM_UTM_sf))
```
Then verify if transformation is successful using this command.

```
> print(crs(ADM_UTM_sf))
CRS arguments:
+proj=utm +zone=47 +datum=WGS84 +units=m +no_defs +ellps=WGS84 +towgs84=0,0,0
```
Get the extents of the PCS and GCS shapefiles. This is needed to calculate the conversion factor (*meter_ reciprocal_PCS2GCS*) from meters to degrees. Compute the conversion factor by getting the lagged differences of xmin and xmax and ymin and ymax for both PCS and GCS. Then compute the ratio of x's and y's of PCS and GCS, add the ratios, and get the average.

```
65 # get boundary box of the shapefile
66 PCS_ext <- extent(ADM_UTM_sf)
67 GCS\_ext \leftarrow extent(ADM\_sf)68
69 # calculate conversion factor from degress to meters using bounding box
70 meter_reciprocal_PCS2GCS <- (diff(PCS_ext[1:2]) / diff(GCS_ext[1:2]) +
71diff(PCS\_ext[3:4]) / diff(GCS\_ext[3:4]))/2
```
STEP 12

Create the grid in three steps:

First, generate an empty raster using **raster()** function through information from GCS extent, degreesconverted-gridsize as the resolution (pixel size) and define the CRS of the blank raster;

```
73 # create an empty raster of grid size granularity
74 ADM_raster <- raster (GCS_ext,
                         res = gridsize/meter_reciprocal_PCS2GCS,
75
                         crs = WGS8476
77
78 # rasterize the shapefile's geocode
    geocode_raster <- fasterize(ADM_sf, ADM_raster, field = "geocode")
79
```
Second, rasterize the shapefile's geocode column. This creates a raster of all the shapefiles' features with the geocodes as raster values.

To get the coordinates of each centroid, convert the raster into dataframe using the function **as.data. frame()** with the option **xy=T** to generate the raster values (geocodes) and its corresponding centroid coordinates.

STEP 14

Use the **head()** command to check the dataframe generated and to learn its structure. The x and y columns are the centroid coordinates. The layer column is the rasterized shapefile attribute (geocode).

STEP 15

Create a new dataframe. Use dplyr's functions and pipe operator (%>%) to perform a series of data manipulations.

First, use **filter()** function to remove all "NA" values in the layer column to get only the centroids inside the country borders.

```
selected.centroids <- geocode_df %>%
87
                                                   # create a new dataframe from geocode_df
      filter(!is.na(layer)) %>%
                                                   # remove NA from the layer columns
88
89
      mutate(id = 1:n()) %>%
                                                   # generate grid ID
90
      select(id,
                                                   # rearrange the columns starting with ID
91
                                                   # rename x centroid coordinate to lon
             1on = x,
92
             lat = y,
                                                   # rename y centroid coordinate to lat
93
             geocode = layer)# lastly, layer column renamed as geocode
```
Second, create a new column containing the grid ID.

Third, rearrange the column starting with ID, x, y, and layer. Rename "x", "y" and "layer" as "lon", "lat", and "geocode", respectively.

```
87
    selected.centroids <- geocode_df %>%
                                                  # create a new dataframe from geocode_df
88
      filter(!is.na(layer)) %>%
                                                  # remove NA from the layer columns
89
      mutate(id = 1:n()) %>%
                                                  # generate grid ID
90
      select(id,
                                                  # rearrange the columns starting with ID
91
             1on = x,# rename x centroid coordinate to lon
             lat = y,92
                                                  # rename y centroid coordinate to lat
93
                                                  # lastly, layer column renamed as geocode
             geocode = layer)
```
STEP 16

Generate the filename for the CSV file output. Indicate the following identifiers:

- country refers to country code;
- "centroid" refers to data content; and
- gridsize and "grid"- refer to the grid size.

Save the centroids dataframe as CSV file. Note that the output path will serve as the working directory.

```
102 - # Output CSV----
103 write.csv(selected.centroids,
104
               file = paste0(file_name,".csv"),
105
               row, names = F)
```
The resulting CSV file should contain the grid ID, centroid coordinates (lon, lat), and the geocode.

In the browser address bar, go to Google Drive⁶ [www.drive.google.com.](http://www.drive.google.com) Click \Box New and then click **File upload**.

⁶ Google Drive is a trademark of Google LLC.

After the file is uploaded, locate the CSV file containing the centroid coordinates.

This file is needed for downloading the satellite imagery of each grid.

Repeat the steps using the country level shapefile. This time, upload the folder containing the country shapefile.

This folder is needed for determining the country boundary.

Downloading Satellite Imagery

STEP 1

In the browser address bar, input the Google Colaboratory (or Colab)7 web address [https://colab.research.](https://colab.research.google.com/) [google.com/](https://colab.research.google.com/) and press **Enter**.

Make sure to log in to your Google account. Then click **Upload**.

⁷ Google Colab is a trademark of Google LLC.

Click **Choose File**.

Locate the Jupyter Notebook file from the computer.

Use **Daytime_imagery_batch_download.ipynb**. Click **Open**.

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STEP 3

Click **Connect**.

This will initialize the Colab's environment.

The Jupyter Notebook has two parts:

■ Text cell is the non-executable part containing code descriptions or headers.

Code cell contains the Python commands and it is denoted by square brackets "[]".

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To execute, click on each code cell and click \bullet button at the beginning of each code cell.

The first code cell sets up and mounts the Google Drive. Click on the link.

STEP 4

In the browser, sign in to your Google account.

Click **Allow**.

Click the **Copy** icon $\boxed{\Box}$ to copy the code.

Return to the Colab browser tab. **Paste** the code in the text box. Then press **Enter**.

A status will show the path where Google Drive is mounted.

STEP 5

Ensure that any edits made in the libraries are automatically reloaded and any charts or images displayed are shown in the notebook.

STEP 6

Setup the Google Earth Engine (GEE).8

⁸ Google Earth Engine is a trademark of Google LLC.

Install GEE Python library to the Colab virtual machine.

Initialize the authentication of the GEE account by clicking on the link.

STEP 7

In the browser, sign in to your Google account.

Click **Allow**.

Click the **Copy** icon $\boxed{}$ to copy the code.

Return to the Colab browser tab. **Paste** the code in the text box. Then press **Enter**.

A status will show that the authorization token has been successfully saved.

STEP 8

Load the GEE library into the Python environment and initialize it.

Read the CSV file that contains the grid centroids.

Load the Python Data Analysis Library (Pandas) package that is used for reading external table files and manipulating data. Fetch the link of the CSV file that was previously uploaded to the Google drive and store it in the **centroid_csv_path** variable.

STEP 10

Click the **Files** icon **the Supersy Section**.

Click **gdrive** from the list of folders and expand the file directory tree to find the CSV file location.

Click the vertical ellipsis to show more file options.

Click **Copy path**.

STEP 14

Paste the link on the blank space after the variable *centroid_csv_path* and enclose it in apostrophes.

Then press \bullet to execute the code cell.

Execute the code cell to set the *id* column as the dataframe's row index and check the contents of the first five rows of the CSV file.

STEP 16

Determine the dataframe's row count using the *count()* function, which should be equal to the number of satellite imagery to be downloaded. The output is saved in the variable *imagery_count*.

Install the GeoPandas Python library in the Colab virtual machine. GeoPandas is an open source project that enables working with geospatial data in Python easier.

Load the GeoPandas library into the Python environment and then load the shapefile as *adm0_shp* variable. Display the first five rows of the shapefile's attribute table. To load the shapefile, fetch the link of the country level shapefile that was previously uploaded to Google Drive.

STEP 18

Click **Files** icon to show the **Files section**.

Click **gdrive** from the list of folders and expand the file directory tree to find the folder containing the country level shapefile.

STEP 20

From the folder, select the country level shapefile (ADM0).

Click the vertical ellipsis to show more file options.

STEP 22

Click **Copy path**.

Paste the link on the blank space after the variable **adm0_shp** and enclose it in apostrophes.

STEP 24

Execute the code cell. The output shows the contents of the shapefile's attribute table. Only one row of features is displayed because it is a country level shapefile.

STEP 25

Generate the bounding box polygon. This code will limit the imagery download from GEE to the country boundaries.

First, create a bounding box polygon using the GeoPandas function envelope.

```
bbox poly = adm0 shp.geometry.envelope # get bounding box polygon
   bbox = bbox poly.to json() # convert bounding box polygon to json
```
Second, convert **bbox_poly** to java script object notation (JSON).

STEP 26

Extract bounding box coordinates from the JSON object.

First, convert the JSON object to a dictionary object.

```
\blacktrianglerightbbox dict = eval(bbox)bbox features dict = bbox dict['features'][0]
   bbox_coordinates = bbox_features_dict['geometry']['coordinates']
```
Second, create a subset of the first feature containing the coordinates. *There is only one feature because it is a country level shapefile*.

Third, create a subset of the dictionary to get only the coordinate values of the bounding box.

STEP 27

Convert the bounding box coordinate into a GEE polygon object.

View the composite imagery to check if the temporal filter used will generate a complete imagery, specifically for Sentinel-2 satellite imagery, covering the entire country.

Input the code pertaining to the country of study by typing country = "code". Then set the year of interest.

Use an if-else statement to select which satellite imagery to use based on the year of interest and to define the image resolution and image size of the corresponding satellite. Based on the satellite information, generate the folder name where the imagery will be stored in the Google Drive. Then generate the filename using the same information.

```
If int(year) >= 2015:
      day sat="ST" # sentinel 2 satelliteimq res = "384"img_size = str(int(img,res)*10)else:
      day sat = "LS" # Landsat
      img_res = "256"img_size = str(int(img-res)*15)# Generate output directory ---
    drive_folder = "__", join([["CNN", "IMGB", country, year, day_sat, img_res,"TIF", img_size])# assemble DIMG filename
    \texttt{DIMG} = \texttt{"\_".join}(\texttt{['CNN\_DIMG", country, year, day\_sat, img\_res, img\_size]})
```
Print out the values of the variables to check if the outputs are correct.

STEP 29

Specify the starting date of the coverage of satellite imagery.

```
# Specify coverage date
    # Satellite imagery coverage starting date:
    start MM = "01"start DD = "01"start_date = "-".join([year,start_MM,start_DD])
    print("Coverage start date: "+start_date)
\Gamma Coverage start date: 2015-01-01
```
Then specify the end date. The end date of the temporal imagery filter needs to be adjusted to have a longer temporal coverage in case it fails to generate a complete imagery for the entire country.

STEP 30

Through the GEE Application Programming Interface (API), filter the satellite imagery collection based on the temporal range (i.e., **start_date** and **end_date**) and country boundary (i.e., **bounding_box**). Visualize the imagery to check if the temporal filter yields complete imagery for the entire country.

```
# import folium library
import folium
if day sat == "ST":
 def maskS2clouds(image):
   qa = image.setect('QA60')# Bits 10 and 11 are clouds and cirrus, respectively.
   cloudBitMask = 1 \le 10cirrusBitMask = 1 \leq 11# Both flags should be set to zero, indicating clear conditions.
   mask = qa.bitwiseAnd(cloudBitMask).eq(0).And(qa.bitwiseAnd(cirrusBitMask), eq(0))return image.updateMask(mask).divide(10000)
 rgbVis = {'min': 0.0,'max': 0.3,'bands': ['B4', 'B3', 'B2']}
 # Filter an image collection.
 cloud masked = ee. ImageCollection('COPERNICUS/S2')\
 .filterDate(start date, end date)\
  .filterBounds(bounding_box).filter(ee.Filter.lt('CLOUDY_PIXEL_PERCENTAGE', 60))\
 .map(maskS2clouds)
 # Take median value
 satellite_imagery = cloud_masked.median().visualize(**rgbVis)
else:
 if int(year) < 2013:
    landsat mission = 'LANDSAT/LE07/C01/T1' # Select Landsat 7
    LS_day_sat = "LS7"else:
    landsat_mission = 'LANDSAT/LCO8/C01/T1' # Select Landsat 8LS_day_sat = "LS8"
```
First, import the Folium library in the Python environment. Folium is a Python visualization library for geospatial data.

STEP 32

Using an if-else statement, select the appropriate filter for the satellite to be used. The satellite is selected based on the availability of coverage of the imagery. Landsat 7 covers the period January 1999 to present and Landsat 8 covers April 2013 to present, while Sentinel-2 imagery covers the period June 2015 to present.

```
# import folium library
    import folium
   if day sat == "ST":
     def maskS2clouds(image):
       qa = image.setect('QA60')# Bits 10 and 11 are clouds and cirrus, respectively.
       cloudBitMask = 1 \ll 10cirrusBitMask = 1 \ll 11# Both flags should be set to zero, indicating clear conditions.
       \texttt{mask = qa.bitwiseAnd(cloudBitMask).eq(0).And (qa.bitwiseAnd(cirrusBitMask).eq(0))}return image.updateMask(mask).divide(10000)
     rgbVis = {'min': 0.0, 'max': 0.3, 'bands': ['B4', 'B3', 'B2']}
     # Filter an image collection.
     cloud_masked = ee.ImageCollection('COPERNICUS/S2')\
      .filterDate(start_date, end_date)\
     .filterBounds(bounding_box).filter(ee.Filter.lt('CLOUDY_PIXEL_PERCENTAGE', 60))\
     .map(maskS2clouds)
     # Take median value
     satellite_imagery = cloud_masked.median().visualize(**rgbVis)
   else:
     if int(year) < 2013:
        landsat_mission = 'LANDSAT/LE07/C01/T1' # Select Landsat 7
        LS day sat = "LS7"
     else:
        landsat_mission = 'LANDSAT/LC08/C01/T1' # Select Landsat 8
        LS_day_sat = "LS8"# Landsat 7 and 8 imageries are available starting January 1999 and April 2013 to present
     filtered_shp = ee.ImageCollection(landsat_mission)\
     .filterDate(start_date, end_date)\
     .filterBounds(bounding_box)
     # Use inbuilt Earth Engine function to create big composite image from the Landsat tiles
     composite = ee.Algorithms.Landsat.simpleComposite(filtered_shp).float();
      # Pansharpening
     **************************
     if LS_day_sat == "LS7";
       rgb = composite.sleepct('B3', 'B2', 'B1').unitScale(0, 255)# For information on Landsat 7 bands,
        # please visit https://developers.google.com/earth-engine/datasets/catalog/LANDSAT_LE07 C01 T1 SR#bands
     if LS day sat == "LS8":
       rgb = composite.select('B4', 'B3', 'B2').unitScale(0, 255)
        # For information on Landsat 7 bands,
        # please visit https://developers.google.com/earth-engine/datasets/catalog/LANDSAT_LC08_C01_T1_SR#bands
     gray = composite.select('B8').unitScale(0, 155)
      # Convert to HSV, swap in the pan band, and convert back to RGB.
     huesat = rgb.rgbToHsv().select('hue', 'saturation')
     satellite_imagery = ee.Image.cat(huesat, gray).hsvToRgb()
```
Filter the imagery collection in GEE. If the basis is the reference year of the study, then employ Sentinel-2. Define the function **maskS2clouds( )**. Using the Sentinel-2 QA60 band, create a cloud mask to filter over the imagery within the temporal range.

STEP 33

rgbVis defines the visualization parameters to be used in the filter.

- Min and max indicate the values to map red, green, and blue (RGB) 8-bit value to 0 and 255, respectively.
- Bands indicate the satellite bands to visualize.
	- B4 refers to red band.
	- B3 refers to green band.
	- B2 refers to blue band.

Apply filter to the **ImageCollection** (i.e., Sentinel 2, or COPERNICUS/S2 as used in this illustration).

- **filterDate()** defines the temporal coverage.
- **filterBounds()** uses the bounding box to limit the filter to the country boundaries.
- **filter(ee.Filter.lt('CLOUDY_PIXEL_PERCENTAGE', 60))** provides the filter to exclude images with more than 60% cloud cover.
- **map(maskS2clouds)** uses the function for creating cloud mask.

```
\int if day_sat == "ST":
     def maskS2clouds(image):
       qa = image.setlect('QA60')# Bits 10 and 11 are clouds and cirrus, respectively.
       cloudBitMask = 1 \ll 10cirrusBitMask = 1 \ll 11# Both flags should be set to zero, indicating clear conditions.
       mask = qa.bitwiseAnd(cloudBitMask).eq(0).And(qa.bitwiseAnd(cirrusBitMask).eq(0))return image.updateMask(mask).divide(10000)
     rgbVis = {'min': 0.0,'max': 0.3,'bands': ['B4', 'B3', 'B2']}
     # Filter an image collection.
     cloud_masked = ee.ImageCollection('COPERNICUS/S2')\
      .filterDate(start_date, end_date)\
      .filterBounds(bounding box).filter(ee.Filter.lt('CLOUDY PIXEL PERCENTAGE', 60))\
      .map(maskS2clouds)
     # Take median value
     satellite_imagery = cloud_masked.median().visualize(**rgbVis)
```
Generate another object containing the median value of the filtered image collection and apply the visualization parameter.

```
if day_sat == "ST":
Ω
      def maskS2clouds(image):
       qa = image.setlect('QA60')# Bits 10 and 11 are clouds and cirrus, respectively.
       cloudBitMask = 1 \ll 10cirrusBitMask = 1 \ll 11# Both flags should be set to zero, indicating clear conditions.
       mask = qa.bitwiseAnd(cloudBitMask), eq(0).And(qa.bitwiseAnd(cirrusBitMask), eq(0))return image.updateMask(mask).divide(10000)
    rgbVis = {'min': 0.0,'max': 0.3,'bands': ['B4', 'B3', 'B2']}
     # Filter an image collection.
     cloud_masked = ee.ImageCollection('COPERNICUS/S2')\
      .filterDate(start_date, end_date)\
      .filterBounds(bounding_box).filter(ee.Filter.lt('CLOUDY_PIXEL_PERCENTAGE', 60))\
      .map(maskS2clouds)
      # Take median value
      satellite_imagery = cloud_masked.median().visualize(**rgbVis)
```
For Landsat satellite imagery, use Landsat 7 for available imagery prior to 2013 and Landsat 8 for available imagery in 2013 and beyond. Assign the Landsat imagery collection to the variable **landsat_mission**.

LANDSAT/LE07/C01/T1 pertains to Landsat 7 imagery collection in GEE and LANDSAT/LC08/C01/T1 pertains to that of Landsat 8.

Apply filter to the selected Landsat **ImageCollection**.

- **filterDate()** defines the temporal coverage.
- **filterBounds()** uses the bounding box to limit the filter to the country boundaries.

```
Q
  else:
     if int(year) < 2013:
         landsat_mission = 'LANDSAT/LE07/C01/T1' # Select Landsat 7
        LS_day_sat = "LS7"else:
         landsat mission = 'LANDSAT/LC08/C01/T1' # Select Landsat 8
         LS day sat = "LS8"
      # Landsat 7 and 8 imageries are available starting January 1999 and April 2013 to present
     filtered\_shp = ee, ImageCollection(landsat\_mission)\.filterDate(start date, end date)\
      .filterBounds(bounding_box)
      # Use inbuilt Earth Engine function to create big composite image from the Landsat tiles
     composite = ee.Algorithms.Landsat.simpleComposite(filtered_shp).float();
      # Pansharpening
      ***************************
     if LS_day_sat == "LS7":
       rgb = composite.sleepect('B3', 'B2', 'B1').unitscale(0, 255)# For information on Landsat 7 bands,
        # please visit https://developers.google.com/earth-engine/datasets/catalog/LANDSAT LE07 C01 T1 SR#bands
     if LS_day_sat == "LS8":rgb = composite.select('B4', 'B3', 'B2').unitScale(0, 255)
       # For information on Landsat 7 bands,
        # please visit https://developers.google.com/earth-engine/datasets/catalog/LANDSAT LC08 C01 T1 SR#bands
      gray = composite.setect('B8').unitScale(0, 155)# Convert to HSV, swap in the pan band, and convert back to RGB.
      huesat = rgb.rgbToHsv().select('hue', 'saturation')
     satellite_imagery = ee.Image.cat(huesat, gray).hsvToRgb()
```
Generate a composite image for the entire country using the filtered ImageCollection. *This command builds the composite from imagery with less cloud cover.*

Pansharpen the Landsat imagery. This is an intermediate data preparation step undertaken to enhance the resolution of the images. Pansharpening combines high resolution panchromatic images (black and white but sensitive to colors) with lower resolution multispectral band images.

First, select the red, green, and blue (RGB) bands from the composite imagery generated. For Landsat 7, RGB bands are designated as B3, B2 and B1, while Landsat 8's RGB bands are designated as B4, B3 and B2.

Pansharpening ************************** if LS day sat == $"LS7":$ rgb = composite.select('B3', 'B2', 'B1').unitScale(0, 255) # For information on Landsat 7 bands, # please visit https://developers.google.com/earth-engine/datasets/catalog/LANDSAT LE07 C01 Tl SR#bands if LS day sat == $"LSS"$: rgb = composite.select('B4', 'B3', 'B2').unitScale(0, 255) # For information on Landsat 7 bands, # please visit https://developers.google.com/earth-engine/datasets/catalog/LANDSAT LC08 C01 T1 SR#bands $gray = composite.setect('B8') .unitscale(0, 155)$ # Convert to HSV, swap in the pan band, and convert back to RGB. huesat = rgb.rgbToHsv().select('hue', 'saturation') satellite_imagery = ee.Image.cat(huesat, gray).hsvToRgb()

Select the panchromatic band.

Convert the RGB image to Hue Saturation Value (HSV) and select only the hue and saturation bands.

Combine the hue, saturation and the panchromatic bands. Then convert it back into RGB to get the upscaled image.

Determine the x and y coordinates of the bounding box polygon's centroid.

```
#########
# Folium map visualization declarations
# get centroid coordinates of bounding box for map view centering
cen x = bbox poly.centroid.x[0]cen y = bbox poly.centroid.y[0]
# create folium object
map = folium.Map(location=[cen_y, cen_x],
                 zoom start=6,width=1280,height=766,
                 attr=day sat)
```
STEP 41

Create a Folium map object. Use the centroid coordinates of the bounding box to indicate the location to display.

- **zoom_start** defines the initial zoom level of the map.
- **width and height** define the size of the map in pixel units.
- **attr** is the map tile attribution (optional) set to display the name of the satellite used as imagery source.

```
#########
# Folium map visualization declarations
# get centroid coordinates of bounding box for map view centering
cen_x = bbox_poly.centroid.x[0]
cen y = bbox poly.centroid.y[0]
# create folium object
map = folium.Map(location=[cen_y, cen_x],
                 zoom start=6,width=1280,height=766,
                 attr=day sat)
```
Get the **mapID** of the filtered satellite imagery.

```
# get mapID of sat imagery image from GEE
ee image map id = ee. Image(satellite imagery).getMapId()
# add sat imagery to map
folium.raster layers.TileLayer(
   tiles = ee image map id['tile fetcher'].url format,
   attr = 'Google Earth Engine',
   name = 'Daytime Imagery',overlay = True,control = True,).add_to(map)
```
STEP 43

Generate a new map layer to visualize the following parameters:

- **tiles** is the map data source. It uses the mapID to get the URL link of filtered satellite imagery from GEE.
- **attr** is the map tile attribution required if the URL link from Earth Engine is used.
- **name** is the layer name appearing in LayerControl.
- **overlay** is set to **True** to indicate that the imagery will be placed over the Folium default base map.
- **control** is set to **True** so that the layer will be included in the LayerControl.

```
# get mapID of sat imagery image from GEE
ee_image_map_id = ee.Image(satellite_imagery).getMapId()
# add sat_imagery to map
folium.raster_layers.TileLayer(
   tiles = ee_image_map_id['tile_fetcher'].url_format,
   attr = 'Google Earth Engine',
   name = 'Daytime Imagery',overlay = True,control = True,).add to (map)
```
Overlay the bounding box polygon.

```
# add bounding box
folium.GeoJson(
   data = bounding box.getInfo(),name = 'Bounding box',style function=lambda feature: {
        'fillColor': '#FFFFFF00',
        'weight' : 3,'fillOpacity' : 0.5,
        \},
    overlay = True,control = True,).add_to(map)
```
STEP 45

Define the map title for Sentinel and Landsat imagery. Insert a reminder to check if the satellite imagery generated is complete.

```
# add map title
if day sat == "ST":map title = "Sentinel-2 Imagery: Please check if composite image is complete"
else:
  map title = "Landsat Imagery: Please check if composite image is complete"
title html = ''''<h3 align="center" style="font-size:16px"><b>{}</b></h3>
             '''.format(map_title)
map.get_root().html.add_child(folium.Element(title_html))
```
Add the **LayerControl** to the map object. Then instruct Python to display the map.

```
map.get_root().html.add_child(folium.Element(title_html))
# add layer control panel
map.add_child(folium.LayerControl())
# Display the map.
display(map)
```
Below is the output of the map visualization code cell.

As the GEE is limited to only 3000 tasks, it is important to determine the number of tasks in queue to prevent errors.

Use the function **get_queued_tasks( )** to identify the number of "Ready" and "Running" tasks from the GEE task list. This function is necessary to verify if there are fewer than 3000 tasks in queue.

STEP 48

Implement the function **get_queued_task_filenames( )** to obtain the filenames of the "Ready" and "Running" tasks on the GEE task list. This function is necessary to avoid file duplication.

```
def get queued task():
 queued task count = 0for queued_task in ee.batch.Task.list():
    if queued_task.state_in ["READY", "RUNNING"]:
        queued task count += 1
 return queued_task_count
def get queued task filenames():
 print("Fetching queued files")
 task filenames = [ ]for queued task in ee.batch.Task.list():
   if queued_task.state in ["READY", "RUNNING"]:
     print(queued task.state+": "+queued task.status()['description'])
     task_filenames.append(queued_task.status()['description'])
 print('---end fetch---\n')return task filenames
```
Define the function for downloading the satellite imagery.

```
import os
def download_satellite_imagery(sat_imagery):
  next batch size = 10 \# Set the number of new tasks to be added after reaching task limit
  target_count = 3000 - next_batch_size # Threshold before creating new tasks
  task_count = get_queued_task()
  queued_filenames = get_queued_task_filenames()
  print('Number of active tasks: {: }.'.format(task_count))
  for i in range(1, imagery_count):
    imagery_file = DIMG + ' (:06d)' .format(i)imagery_filepath = '/content/gdrive/MyDrive/' + drive_folder + '/' + imagery_file + '.tif'
    if task_count == 3000: # Number of tasks has reached the limit
     # Loop until the task count has not reach the target_count
     while task count > target count:
        active\_task = get\_queued\_task() \# Get the number of tasks on the listif active_task < task_count: # Check if there are finished tasks
         task_count = active_task
          print("***********************")
          print('Number of current pending tasks in queue: {: },' format(task_count))
          print('Remaining tasks before starting new batch: {: }.'.format(task_count-target_count))
          #print(task.status())
    else:
      if (os.path.exists(imagery_filepath) == False):
        if (imagery_file not in queued_filenames):
          print("---------------------")
           print("Starting new task...")
           print("downloading " + imagery_file)
           c_{1}on = df['lon'][i]c_1 at = df['lat'][i]
            geometry = ee.Geometry.Point([c_lon, c_lat]).buffer(1920) #Based imagery resolution of 25
            geometry = geometry.getInfo()['coordinates'][0]
```
First, import the operating system (**os**) library to enable Python to execute operating system commands. In this case, access the folders of the Colab virtual machine.

```
import os
def download_satellite_imagery(sat_imagery):
  next_batch_size = 10 # Set the number of new tasks to be added after reaching task limit
  target_count = 3000 - next_batch_size # Threshold before creating new tasks
  task_count = get_queued_task()
  queued_filenames = get_queued_task_filenames()
  print('Number of active tasks: {: }.'.format(task_count))
```
Define the function *download_satellite_imagery*, which requires a satellite imagery object (**sat_imagery**) as input.

next_batch_size refers to the number of new imagery downloading tasks to be pooled.

target_count refers to the number of tasks in the task list to trigger pooling of new batch of tasks.

Execute the function **get_queued_task()** to determine the number of "Ready" and "Running" tasks in the GEE task list, if any. Then store it in the *task_count* variable. Get the list of "Ready" and "Running" tasks' filenames, if any, by calling the function *get_queued_task( )* and store it in *queued_filenames* variable. Lastly, print out the number of active tasks.

Loop through the list of grid centroids and download the images. The for-loop range is the number of centroids in the CSV file.

```
for i in range(1, imagery_count):
  imagery_file = DIMG + ' '(:06d)' .format(i)\texttt{imagery\_filepath} = \frac{1}{\text{content}/\text{gdrive}/\text{MyDirive}}' + \text{drive\_folder} + \frac{1}{2} + \text{imagery\_file} + \frac{1}{2} + \text{if'}if task count == 3000: # Number of tasks has reached the limit
    # Loop until the task count has not reach the target_count
    while task_count > target_count:
      active\_task = get\_queued\_task() \# Get the number of tasks on the listif active_task < task_count: # Check if there are finished tasks
        task_count = active\_taskprint("***********************")
       print('Number of current pending tasks in queue: {: }.'.format(task_count))
        print('Remaining tasks before starting new batch: {: }, '.format(task_count-target_count))
```
Declare the imagery filename (**imagery_file**) to be used and its complete file path (**imagery_filepath**).

```
for i in range(1, imagery_count):
 imagery_file = DIMG + ' '(:06d)' .format(i)imagery_filepath = '/content/gdrive/MyDirive/' + drive_folder + '/' + imagery_file + '.ti'if task_count == 3000: # Number of tasks has reached the limit
   # Loop until the task count has not reach the target_count
   while task_count > target_count:
     active\_task = get\_queued\_task() \# Get the number of tasks on the listif active_task < task_count: # Check if there are finished tasks
       task_count = active_task
       print("***********************")
       print('Number of current pending tasks in queue: {: }, '.format(task_count))
       print('Remaining tasks before starting new batch: {: }.'.format(task_count-target_count))
```
Implement an if-statement to limit the number of tasks in queue and to prevent errors. *If the task_count reaches 3000, it stops creating new tasks.*

The while-loop will check if the current **task_count** has reached the set threshold (**target_count**) before creating a new batch of tasks

```
for i in range(1, imagery count):
 imagery file = DIMG + ' {:06d}'.format(i)
 \frac{1}{1} imagery_filepath = '/content/gdrive/MyDrive/' + drive_folder + '/' + imagery_file + '.tif'
  if task_count == 3000: # Number of tasks has reached the limit
   # Loop until the task count has not reach the target count
   while task_count > target_count:
     active\_task = get\_queued\_task() \# Get the number of tasks on the listif active_task < task_count: # Check if there are finished tasks
       task_count = active\_taskprint("***********************")
       print('Number of current pending tasks in queue: {: }, '.format(task_count))
       print('Remaining tasks before starting new batch: {: }.'.format(task_count-target_count))
```
The if-statement checks for finished tasks and prints out information on the number of tasks currently in queue and a countdown of when a new batch of tasks will be created.

```
for i in range(1, imagery count):
 imagery_file = DIMG + ' {'.06d}' format(i)imagery_filepath = '/content/gdrive/MyDirive/' + drive_folder + '/' + imagery_file + '.ti'if task count == 3000: # Number of tasks has reached the limit
   # Loop until the task count has not reach the target count
   while task count > target count:
     active\_task = get\_queued\_task() \# Get the number of tasks on the listif active_task < task_count: # Check if there are finished tasks
       task_count = active_task
       print("***********************")
       print('Number of current pending tasks in queue: {: }.'.format(task_count))
       print('Remaining tasks before starting new batch: {: }.'.format(task_count-target_count))
```
If the number of tasks is fewer than 3000 or if the new batch of tasks needs to be created, first check whether the new imagery to be pooled is already in the Google Drive or in queue. *This verification will prevent duplication of tasks.*

```
O
        else:
          if (os.path.exists(imagery filepath) == False):
            if (imagery file not in queued filenames):
                print("----------------------")
                print("Starting new task...")
                 print("downloading " + imagery_file)
                c_lon = df['lon'][i]c_lat = df['lat'][i]geometry = ee.Geometry.Point([c_lon, c_lat]).buffer(1920) #Based image:
                 geometry = geometry.getInfo()['coordinates'][0]
                if (\text{day\_sat} == "ST"):
                   scale = 10elif (\text{day}\_\text{sat} == "LS"):
                   scale = 15task_config = {
                     'scale': scale,
                     'region': geometry,
                    'driveFolder': drive_folder,
                 \mathcal{F}task = ee.batch.Export.image(sat_imagery, imagery_file, task_config)
                 task.start()
                task_count += 1if task_count % 1000 == 0:
                  task_count = get_queued_task()
                print('Number of active tasks: {: }.'.format(task_count))
            else:
                print("On queue: " + imagery_file + ".tif")
          else:
            print("Downloaded: " + imagery_file + ".tif")
```
Print to determine whether the files are in the save path or if they are still in queue.

```
\bulletelse:
          if (os.path.exists(imagery filepath) == False):
            if (imagery_file not in queued_filenames):
                 print("----------------------")
                 print("Starting new task...")
                 print("downloading " + imagery_file)
                 c_{lon} = df['lon'][i]c lat = df['lat'][i]geometry = ee. Geometry. Point([c lon, c lat]). buffer(1920) #Based image:
                 geometry = geometry.getInfo()['coordinates'][0]
                 if (\text{day}\_\text{sat} == "ST"):
                  scale = 10elif (\text{day}\_\text{sat} == "LS"):
                  scale = 15task\_config = {'scale': scale,
                     'region': geometry,
                     'driveFolder': drive_folder,
                 \mathbf{r}task = ee.batch.Export.image(sat_imagery, imagery_file, task_config)
                 task.start()
                task_count += 1if task_count % 1000 == 0:
                   task_count = get_queued_task()
                 print('Number of active tasks: {: }.'.format(task_count))
            else:
                 print("On queue: " + imagery_file + ".tif")
          else:
            print("Downloaded: " + imagery_file + ".tif")
```
Set **c_lon** and **c_lat** (i.e., longitude and latitude, respectively) to store the centroid coordinates obtained from the centroid CSV.

```
Q
        else:
          if (os.path.exists(imagery_filepath) == False):
            if (imagery file not in queued filenames):
                print("---------------------")
                print("Starting new task...")
                print("downloading " + imagery file)
                c lon = df['lon'][i]c lat = df['lat'][i]geometry = ee.Geometry.Point([c_lon, c_lat]).buffer(1920) #Based image:
                geometry = geometry.getInfo()['coordinates'][0]
                if (day sat == "ST"):
                  scale = 10elif (\text{day}\_\text{sat} == "LS"):
                  scale = 15task config = {
                    'scale': scale,
                    'region': geometry,
                    'driveFolder': drive_folder,
                \mathbf{r}task = ee.batch.Export.image(sat imagery, imagery file, task config)
                task.start()
                task count += 1if task_count % 1000 == 0:
                  task_count = get_queued_task()
                print('Number of active tasks: {: }.'.format(task count))
            else:
                print("On queue: " + imagery_file + ".tif")
          else:
            print("Downloaded: " + imagery_file + ".tif")
```
Employ the centroid coordinates to define a geospatial circle using a GEE point geometry with a buffer of 1920 meters. *This buffer value corresponds to half of the grid size measured from the centroid to the grid boundary*.

As illustrated in Step 6 of the section on Generating Centroids for Satellite imagery, buffer size is computed as follows:

256 pixel x 15 meters/pixel = 3840 meter grid size 3840 / 2 = 1920 meter buffer size

where: 15 meters/pixel is the Landsat resolution

```
\mathbf 0else:
          if (os.path.exists(imagery filepath) == False):
            if (imagery file not in queued filenames):
                print("----------" )print("Starting new task...")
                print("downloading " + imagery file)
                c_lon = df['lon'][i]c_1lat = df['lat'][i]
                 geometry = ee. Geometry. Point([c lon, c lat]). buffer(1920) #Based image:
                 geometry = geometry.getInfo()['coordinates'][0]
                if day_sat == "ST"):
                   scale = 10elif (\text{day}\text{-} \text{sat} == "LS"):
                 scale = 15task_config = {
                     'scale': scale,
                     'region': geometry,
                     'driveFolder': drive folder,
                \mathcal{F}task = ee.batch.Export.image(sat_imagery, imagery_file, task_config)
                task.start()
                task_count += 1if task count \frac{1000}{1000} = 0:
                  task count = get queued task()print('Number of active tasks: {: }.'.format(task_count))
            else:
                print("On queue: " + imagery_file + ".tif")
          else:
            print("Downloaded: " + imagery_file + ".tif")
```
Redefine the geometry variable using the coordinates of the circle as its value.

```
\bulletelse:
          if (os.path.exists(imagery filepath) == False):
             if (imagery_file not in queued_filenames):
                print("----------------------")
                print("Starting new task...")
                 print("downloading " + imagery_file)
                c_{10n} = df['lon'][i]c<sub>lat</sub> = df['lat'][i]
                 geometry = ee.Geometry.Point([c_lon, c_lat]).buffer(1920) #Based imager
                 geometry = geometry.getInfo()['coordinates'][0]
                 if (day sat == "ST"):
                   scale = 10elif (\text{day} \text{ sat} == "LS");scale = 15task_config = {
                     'scale': scale,
                     'region': geometry,
                     'driveFolder': drive folder,
                 \mathcal{F}task = ee.batch.Export.image(sat_imagery, imagery_file, task_config)
                 task.start()
                 task_count += 1if task_count \frac{1000}{1000} = 0:
                   task count = get queued task()print('Number of active tasks: {: }.'.format(task_count))
            else:
                 print("On queue: " + imaginary file + ".diff")else:
            print("Downloaded: " + imagery_file + ".tif")
```
Next, define the export parameter using the task_config dictionary variable.

The **task_config** is composed of the following:

- **scale** is the satellite resolution (10 meter/pixel Sentinel; 15 meter/pixel Landsat),
- **region** is the area coverage to download, and
- **driveFolder** is the folder path where the downloaded imagery will be stored.

```
\mathbf \Omegaelse:
          if (os.path.exists(imagery filepath) == False):
             if (imagery_file not in queued_filenames):
                 print("----------" )print("Starting new task...")
                 print("downloading " + imagery file)
                c_{\text{lon}} = df['lon'][i]c<sub>lat</sub> = df['lat'][i]
                 geometry = ee. Geometry. Point([c lon, c lat]). buffer(1920) #Based image:
                 geometry = geometry.getInfo()['coordinates'][0]
                 if (\text{day sat} == "ST"):
                   scale = 10elif (\text{day\_sat} == "LS"):
                  scale = 15task\_config = {'scale': scale,
                     'region': geometry,
                     'driveFolder': drive_folder,
                 \mathcal{F}task = ee.batch.Export.image(sat_imagery, imagery_file, task_config)
                 task.start()
                 task_count += 1if task count % 1000 == 0:
                  task_count = get_queued_task()
                 print('Number of active tasks: {: }.'.format(task_count))
             else:
                 print("On queue: " + imagery_file + ".tif")
          else:
             print("Downloaded: " + imagery_file + ".tif")
```
Describe the image batch export object and name it as task. The image batch export object requires the following parameters:

- satellite imagery (**sat_imagery**),
- **F** filename to be used (**imagery_file**), and
- **E** export parameter (**task_config**).

Finally, pass the task to GEE using the command **task.start()** and add another task to the task counter variable **task_count**.

```
\bulletelse:
          if (os.path.exists(imagery filepath) == False):
             if (imagery_file not in queued_filenames):
                 print("---------------------")
                 print("Starting new task...")
                 print("downloading " + imagery_file)
                 c_lon = df['lon'][i]c_lat = df['lat'][i]geometry = ee. Geometry. Point([c lon, c lat]). buffer(1920) #Based image:
                 geometry = geometry.getInfo()['coordinates'][0]
                 if (\text{day sat} == "ST"):
                   scale = 10elif (\text{day}\_\text{sat} == "LS"):
                   scale = 15task_config = {
                     'scale': scale,
                     'region': geometry,
                     'driveFolder': drive folder,
                 \mathcal{F}task = ee.batch.Export.image(sat_imagery, imagery_file, task_config)
                 task.start()
                 task count += 1if task count \frac{1000}{1000} = 0:
                   task_count = get_queued_task()
                 print('Number of active tasks: {: }.'.format(task_count))
             else:
                 print("On queue: " + imaginary file + ".diff")else:
             print("Downloaded: " + imagery_file + ".tif")
```
Provide printouts of the number of tasks being pooled. To speed up the task creation process, execute **get_queued_task()** only after every 1000 tasks to check the exact number of tasks in queue.

```
\mathbf 0else:
          if (os.path.exists(imagery_filepath) == False):
            if (imagery_file not in queued_filenames):
                print("----------------------")
                print("Starting new task...")
                print("downloading " + imagery_file)
                c_lon = df['lon'][i]c lat = df['lat'][i]geometry = ee.Geometry.Point([c_lon, c_lat]).buffer(1920) #Based image:
                 geometry = geometry.getInfo()['coordinates'][0]
                if (day sat == "ST"):
                   scale = 10elif (\text{day}\text{-} \text{sat} == "LS");scale = 15task config = {
                     'scale': scale,
                     'region': geometry,
                     'driveFolder': drive_folder,
                 \mathbf{H}task = ee.batch.Export.image(sat_imagery, imagery_file, task_config)
                task.start()
                task count += 1if task_count % 1000 == 0:
                   task_count = get_queued_task()
                print('Number of active tasks: {: }.'.format(task_count))
            else:
                print("On queue: " + imagery_file + ".tif")
          else:
            print("Downloaded: " + imagery_file + ".tif")
```
Implement the function **download_satellite_imagery()** and pass it on to the filtered GEE imagery stored in the object **satellite_imagery** as the function's argument. *As the function runs, it prints out the task information.*

The following is the function printout when restarting the imagery download process, which displays all the files that are still in queue.

Below is the printout of the number of pending tasks and the downloaded and pending imagery, which were skipped to avoid duplication.

```
- Download satellite imagery
   o download_satellite_imagery(satellite_imagery)
    <sup>2</sup> Fetching queued files
        READY: CNN DIMG THA 2015 ST 384 3840 000009
        RUNNING: CNN_DING_THA_2015_ST_384_3840_000008
        RUNNING: CNN DIMG THA 2015 ST 384 3840 000007
        RUNNING: CNN_DIMG_THA_2015_ST_384_3840_000006
        RUNNING: CNN_DIMG_THA_2015_ST_384_3840_000005
        RUNNING: CNN_DIMG_THA_2015_ST_384_3840_000004
         --- end fetch---
        Number of active tasks: 7.
        \begin{tabular}{ll} \textbf{Downloaded:} & \textbf{CNN\_DING\_THA\_2015\_ST\_384\_3840\_000001.tif}\\ \textbf{Downloaded:} & \textbf{CNN\_DING\_THA\_2015\_ST\_384\_3840\_000002.tif}\\ \end{tabular}Downloaded: CNN_DING_THA_2015_ST_384_3840_000003.trlf<br>Downloaded: CNN_DING_THA_2015_ST_384_3840_000004.trlf<br>Downloaded: CNN_DING_THA_2015_ST_384_3840_000005.trlf
        On queue: CNN_DIMG_THA_2015_ST_384_3840_000006.tif
        On queue: CNN DIMG THA 2015 ST 384 3840 000007.tif
        On queue: CNN_DIMG_THA_2015_ST_384_3840_000008.tif
        On queue: CNN_DIMG_THA_2015_ST_384_3840_000009.tif
        Starting new task...
        downloading CNN_DIMG_THA_2015_ST_384_3840_000010
        Number of active tasks: 8.
        Starting new task...
        downloading CNN_DIMG_THA_2015_ST_384_3840_000011
        Number of active tasks: 9.
        Starting new task...
        downloading CNN_DIMG_THA_2015_ST_384_3840_000012
        Number of active tasks: 10.
```
Saving of imagery from the GEE to Google Drive consumes some time. Depending on the quantity of imagery to download, the 12-hour Colab runtime may not suffice. Thus, it is necessary to run re-run the code. In the browser, go back to Google Drive and verify if the files are downloaded.

Click the folder name to verify if the files are downloaded.

Download all images for specific country and year. Click the folder name to reveal folder options. Then press **Download**.

The download process starts after Google Drive has finished compressing the files.

Save the ZIP file in the working folder and then **unzip the file**.

Converting Format of Satellite Imagery

Use the Geospatial Data Abstraction Library (GDAL) to convert images into geo-tagged image file format (geoTIFF). Crop the images to get the correct number of pixels. Prepare a *.tar.gz archive file of all input JPG images for easier handling in Colab.

Use the R code: *Daytime_imagery_format_conversion.R*.

Load the **tidyverse** and **gdalUtilities** packages.

Select the working directory using the function **tk_choose.dir()** from the package **tcltk**. This function opens a window for choosing the directory containing the daytime satellite imagery. Set the folder path to **sat_imagery_folder**.

Using the **setwd()** command, set the previously assigned folder (i.e., **sat_imagery_folder** in this illustration) as the working directory.

```
11 - # Select location of satellite imagery ---
12 # NOTE: Make sure to double click the folder to make the selection
13 sat_imagery_folder <- tcltk::tk_choose.dir(caption = "Select Directory Containing Daytime Satellite Imagery")
1415 # set working directory using the directory holding satellite imagery folder
(16 setwd(dirname(sat_imagery_folder))
```
Use the function **tk_choose.dir()** from the package **tcltk** to open a window to select the CSV file containing the grid centroids used to download the satellite imagery.

```
18 # Load the centroid csv--
19 # select the csv path from the open dialog window
20 csv_path <- tcltk::tk_choose.files(filters = matrix(c("CSV",".csv","All files","*"),2,2,byrow = T),
21caption = "Select Grid Centroid CSV")
2223 # read centroid csv
24 df_centroid <- read.csv(csv_path, stringsAsFactors = F)
```
Load the CSV file as a **df_centroid** dataframe.

```
18 # Load the centroid csv---
19 # select the csv path from the open dialog window
20 csv_path <- tcltk::tk_choose.files(filters = matrix(c("CSV",".csv","All files","*"),2,2,byrow = T),
21
                                       caption = "Select Grid Centroid CSV")
22
23 # read centroid csv
24 df_centroid <- read.csv(csv_path, stringsAsFactors = F)
```
STEP 4

Create a destination folder using the function **str_replace( )** to change the character **TIF** from the variable **sat_imagery_folder** into **JPG**.

```
26+ # Check if destination folders exists, otherwise create folders----
27 # create destination folder name
28 dest_path <- paste0("./",str_replace(basename(sat_imagery_folder),"_TIF_","_JPG_"))
29
30 - if (!dir.exists(dest_path)) {
31dir.create(dest_path)
32 - 3
```
STEP 5

Create a new dataframe from **df_centroid**. In this dataframe, generate two columns containing the full path of the TIF and JPG filenames and a separate column containing only the filename of the JPG files without the file path.

```
34 - # Create new columns to hold the tif and jpg file paths ---
35
36 df <- df_centroid %>%
   mutate(tif_file=list.files(sat_imagery_folder,pattern = ".tif$",full.names = T)) %>%
37
38
   mutate(jpg_file=paste0(dest_path,"/",
39
                             str_replace(basename(tif_file),".tif",".jpg"))) %>%
40
     mutate(filename=basename(jpg_file))
```
Set the pixel resolution of each imagery based on the source satellite. Using the function **str_detect()**, check the satellite imagery folder name for the embedded satellite code name.

```
42 + # Detect satellite imagery source embedded on folder name ---
43 - if (str_detect(sat_imagery_folder,"ST")) {
      img\_res = 38444
45 + }else if (str_detect(sat_imagery_folder,"LS")) {
46
      img\_res = 25647 - 3
```
STEP 7

Define the function to crop and convert the TIF files into JPG files. *It takes the filename and path* of the TIF and JPG files as input. The function also prints out the TIF and JPG filename that are being processed.

STEP 8

Employ the function **gdal_translate( )** from the gdalUtilities package to execute this task through the following parameters:

- **src_dataset** is the file path of the TIF input file.
- **dst_dataset** is the file path of the JPG output file.
- **srcwin = c(xoff,yoff,xsize,ysize)** selects a sub window from the source image for copying based on pixel/line location and specify pixel count based on the satellite imagery source.
- **of** refers to the output format "JPEG".
- **scale** is set to "" so that the input pixel values will not be changed.
- **co** passes a creation option to the output format driver. This sets the JPEG output quality to 100% or no compression.

Implement *apply()* function to go through each row of the TIF file listed in the dataframe and pass it on to the custom function *process_imagery()*.

STEP 10

Remove the column containing TIF and JPG file path.

```
64 df \leq - df \frac{9}{2}select(-c(tif_file,jpg_file))
65
```
Create a vector shapefile using the centroids coordinates. Load the package **sf**. Define the Coordinate Reference System (CRS) variable for the shapefile.

STEP 12

Generate a duplicate of the centroid coordinates to preserve the data inside the shapefile's attributes. Then using the sf function *st_as_sf()*, create the shapefile. *This will be used later in aggregating luminosity values in GEE*.

```
67 * # Create multipoint vector shapefiles from the dataframe----
68 # This shall be use later in aggregating luminosity values in GEE
69
70 library(sf)
7172 - # Define crs variable ----
73 WGS84 <- "+proj=longlat +datum=WGS84 +no_defs +ellps=WGS84 +towgs84=0,0,0"
74
   pt\_shp \leftarrow df \gg\%75
      mutate(x = lon, y = lat) %76
      st\_{as\_{sf}}sf(coords = c("x","y"), crs = WGS84)
77
```
Generate the filename for the shapefile. Prefix the centroid's CSV filename with "shp" and change the file extension to ".shp". Then output the vector shapefile. *The shapefile is needed in for aggregating luminosity values of each grid in GEE in the subsequent steps.*

STEP 14

Create a gzip (.tar.gz) archive file containing the JPG files.

First, specify the filename of the archive file. Then use the **tar( )** function to compress the JPG folder through the following parameters:

- **tarfile** is the output filename,
- **file** is the destination path, and
- **compression** is the archive file type "gzip".

```
92 - # Create tar.gz archive file----
93 tar_filename <- paste0(sub("^.+/","", dest_path),".tar.gz")
94
95 tar(tarfile = tar_filename,96
        files = dest_path,97
        compression = "gzip")
```
The JPG output folder and tar.gz file are saved in the same folder as the TIF folder.

Then click **File upload**.

Locate and select the **tar.gz** archive file containing the JPG images.

Nighttime Satellite Imagery Processing

The following sections detail how to download nighttime satellite imagery and aggregating luminosity values.

Nighttime lights (NTL) imageries covering 1992 to 2013 are available from the Defense Meteorological Program (DMSP) Operational Line-Scan System (OLS) while NTL imageries covering 2012 to 2020 are available from the Visible Infrared Imaging Radiometer Suite (VIIRS). DMSP-OLS and VIIRS imagery are both hosted by the Earth Observation Group, Colorado School of Mines.

DMSP-OLS data are available as global coverage per year per image and can be downloaded from this link: *<https://eogdata.mines.edu/dmsp/downloadV4composites.html>*.

VIIRS imagery are published as daily mosaic and monthly and annual composite images. Unlike DMPS-OLS, VIIRS imagery is split into 6 tiles. Information on VIIRS NTL version 1 data is available from this link: https://eogdata.mines.edu/products/vnl/. When downloading, take note of the tile where the country of interest is covered.

For VIIRS, only 2015 and 2016 have annual composite images. Thus, GEE is used to create an annual composite for years other than those aforementioned using the monthly composite imagery.

STEP 1

Download VIIRS nightlight satellite imagery version 1 for years with available annual composite images.

In the browser, go to the VIIRS website https://eogdata.mines.edu/nighttime_light/annual/v10/. **Select**, and **click** the required year (e.g., "**2015**").

Select the tile **where the country of interest is located**. The tile information is the fourth group of characters from the right. **Save** the file in the working directory. Note that the file is a tar.gz archive with a size of approximately 4 GB.

Once download has finished, **decompress** the archive file.

Crop out the nighttime imagery for the country of interest.

Open the Rcode **Crop_NTL_imagery.R** in Rstudio. From the top right bars, click **the Source** \cdot to run the entire script.

Load the required packages.

Use **tk_choose.files()** from the package **tcltk** to open a window for selecting and obtaining the country level shapefile path. *Please note that country level shapefiles are usually denoted as ADM0*.

```
8+ # Opens a dialog box for selecting country shapefile --
9 shapefile_path <- tcltk::tk_choose.files(filters = matrix(c("SHP",".shp","All files","*"),2,2,byrow = T)
10
                                            caption = "Select Country Level Shapefile")
11 # read shapefile
12 shp <- read_sf(shapefile_path)
13
14 # extract bounding box and round up the values to add some buffer
15 xmin <- floor(st_bbox(shp)[[1]])
16 ymin <- floor(st_bbox(shp)[[2]])
17 xmax <- ceiling(st_bbox(shp)[[3]])
18 ymax <- ceiling(st_bbox(shp)[[4]])
```


Load the shapefile using the sf function **read_sf( )**.

```
8 + # Opens a dialog box for selecting country shapefile----
 9 shapefile_path <- tcltk::tk_choose.files(filters = matrix(c("SHP",".shp","All files","*"),2,2,byrow = T),
10
                                            caption = "Select Country Level Shapefile")
11 # read shapefile
12 shp <- read_sf(shapefile_path)
13
14 # extract bounding box and round up the values to add some buffer
15 xmin <- floor(st_bbox(shp)[[1]])16 ymin <- floor(st_bbox(shp)[[2]])
17 xmax <- ceiling(st_bbox(shp)[[3]])
18 ymax <- ceiling(st_bbox(shp)[[4]])
```
STEP 6

Extract the bounding box coordinates of the shapefile using the function *st_bbox()* from the *sf* package. Expand the bounding box to have some buffer. *This can be done by rounding down ymin and xmin, and rounding up ymax and xmax*.

```
8 + # Opens a dialog box for selecting country shapefile----
 9 shapefile_path <- tcltk::tk_choose.files(filters = matrix(c("SHP",".shp","All files","*"),2,2,byrow = T),
                                            caption = "Select Country Level Shapefile")
10
11 # read shapefile
12 shp <- read_sf(shapefile_path)
13
14 # extract bounding box and round up the values to add some buffer
15 xmin <- floor(st_bbox(shp)[[1]])
16 ymin <- floor(st_bbox(shp)[[2]])
17 xmax <- ceiling(st_bbox(shp)[[3]])
18 ymax <- ceiling(st_bbox(shp)[[4]])
```
STEP 7

Select the directory containing the nighttime satellite imagery using the function *tk_choose.dir()* from the package *tcltk*.

```
20 - # Opens a dialog box for selecting geoTIFF NTL data---
21 NTL_file_folder <- tcltk::tk_choose.dir(caption = "Select Directory Containing Nighttime Satellite Imagery")
22+ # Get working directory from downloaded NTL data-
23 wd_path <- dirname(NTL_file_folder)
24 setwd(wd_path)
```


A window opens for selecting the directory containing the nighttime satellite imagery.

STEP 8

Extract the parent folder path of the nighttime satellite imagery and use it as the working directory through the **setwd()** command.

Obtain the filenames of all nighttime satellite imagery files that are stored in the folder.

STEP 10

Use an if-else statement to select the correct imagery product.

- For VIIRS, use the data product **vcm-orm-ntl** with extension **avg_rade9.tif**.
- For DMSP-OLS, use data product **web.stable_lights.avg_vis**.

```
26 + # Check if correct file is selected reselect if needed----
27 # Filter NTL data products:
28 # for VIIRS: vcm-orm-ntl with extension avg_rade9.tif
29 # for DMSP: web.stable_lights.avg_vis
30 NTL_file_list <- list.files(path = NTL_file_folder,
31pattern = ".tifs",32full.name = T)33
34 - if (str_detect(NTL_file_folder,"SVDNB_npp")) {
35<sup>°</sup># filter for VIIRS
36 NTL_file <- NTL_file_list[str_detect(NTL_file_list,"vcm-orm-ntl")]
37 - \} else{
38
      # filter for DMPS
39
      NTL_file <- NTL_file_list[str_detect(NTL_file_list,"web.stable_lights.avg_vis")]
40 - 741
42 print(basename(NTL_file))
```
Print the filename to check.

```
> print(basename(NTL_file))
[1] "SVDNB_npp_20150101-20151231_75N060E_vcm-orm-ntl_v10_c201701311200.avg_rade9.tif"
×
```
Generate the destination path where the cropped nighttime imagery and base name for the output file will be saved.

```
44 + # Generate destination folder and output file ----
45 dest_path <- paste0(wd_path,"/cropped_",basename(NTL_file_folder),"/")
46 output_file <- paste0(dest_path,"cropped_",basename(NTL_file))
4748 - # Check if destination folders exists, otherwise create folders ----
49 - if (!dir.exists(dest_path)) {
50 dir.create(dest_path)
51 - \}
```
STEP 12

Check if the destination folder already exists. If the folder does not exist yet, create it.

```
44 + # Generate destination folder and output file ----
45 dest_path <- paste0(wd_path,"/cropped_",basename(NTL_file_folder),"/")
46 output_file <- paste0(dest_path,"cropped_",basename(NTL_file))
47
48 * # Check if destination folders exists, otherwise create folders ----
49 - if (!dir.exists(dest_path)) {
    dir.create(dest_path)
50
51 - 3
```
STEP 13

Run the *gdal_translate( )* function from the *gdalUtilities* package to crop the nighttime satellite imagery.

```
53 - # Crop the NTL image----
54
55 gdal_translate(NTL_file,output_file,projwin = c(xmin,ymax,xmax,ymin))
```
STEP 14

The code's output is stored in the folder with a prefix "cropped_". Likewise, the geoTIFF file is prefixed. It will later be uploaded to GEE for further processing.

Compute the aggregate average luminosity per area, where every pixel's night light intensity is considered.

Aggregation computation is done in GEE, where the shape for each area needs to be defined and nighttime imagery for corresponding year needs to be provided. The total sum is divided by the number of pixels.

Use the code in file: *viirs_mean_luminosity.js*.

Upload the cropped nighttime lights imagery. Click **Assets.**

New Script - Earth Engine Cod X + \rightarrow C \bullet code.earthengine.google.com \star 0 $Q \rightarrow$ Google Earth Engine Search places and datasets. ◎ 国 皇 Scripts Doct Assets **Run - Reset - Apps Co. Inspector Console** Tasks New Scrip Get Link Use print(...) to write to this console. $W - C$ 1 Filter scripts. A Dwner (2)
A Writer
A Reader
A Archive
No accession
again. Attention Python and JavaScript client
1ibrary users!
Earth Engine servers now require client
1ibrary v0.1.215, released March 11.
Please update to the latest Fython or
2iavaScript version to avoid a break in
service. ible repositories. Click Refresh to check O V V V II Map Satellite \mathbb{S}^2 $\frac{+}{-}$ North
Pacific
Ocean

STEP 16

Click **New**.

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STEP 17

Click **GeoTIFF**.

STEP 18

Click **Select** and locate the cropped nighttime lights imagery.

Change the **Asset ID**. Make sure that the ID only contains letters and numbers.

STEP 20

Click **Upload**.

The uploaded nighttime lights data will appear as a new asset.

STEP 21

This time upload the point shapefile. Again click **New** and select **Shape files**.

Click **Select**.

STEP 23

Locate the shapefile that was created from the code *Daytime_imagery_format_conversion.R*.

Click **Upload**.

The uploaded shapefile will appear as a new asset.

Open the JavaScript **ntl_mean_luminosity.js** using a text editing software (e.g., Windows Notepad).

```
...
                                                                                 O] ntl_mean_luminosity.js
var annual_composite = viirs_annual<br>select('b1')//Average DNB radiance values
÷.
var nlVis = \ellmin: 0.0,<br>max: 1,<br>bands: ['b1'],
\mathbf{H}Map.centerObject(<u>pt_shp</u>);<br>Map.addLayer(<u>annual_composite.nlVis</u>,"NTL annual composite");
// Aggregate mean of nightlight intensities for centroid regions of size 256px
var mappedFeatures = pt_shp.map(function(feature) {<br>var geometry = ee.Geometry.Point(lee.Number(et_shp.get('lon')), ee.Number(et_shp.get('lat')))).buffer(1920).bounds()<br>return feature.set(annual_composite.reduceRegion({
      reducer: 'mean',<br>geometry: feature.geometry(),
       scale: 100,
v_{\rm{in}}// Export the FeatureCollection.<br>Export.table.toDrive({
   collection: mappedFeatures,<br>description: '',<br>fileFormat: 'CSV'
);
```
Select "b1" band of viirs_annual raster and store it in the variable **annual_composite**.

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Define variable **nlVis** to store the map visualization parameters.

Use the grid centroid shapefile, which will be imported later, to put the map view in the center.

Visualize b1 band of the viirs_annual raster using visualization parameters defined in nlVis through the command **Map.addLayer( )**.

```
0.0.0<sup>2</sup> ntl_mean_luminosity.js
var annual_composite = viirs_annual<br>select('b1')//Average DNB radiance values
÷.
var nlVis = \ellmin: 0.0,<br>max: 1,<br>bands: ['b1'],
}:
Map.centerObject(<u>ot_sho);</u><br>Map.addLayer(<u>annual_composite.nlVis</u>,"NTL annual composite");
// Aggregate mean of nightlight intensities for centroid regions of size 256px
var mappedFeatures = pt_shp.map(function(feature) {<br>var geometry = <u>ee.Geometry.Point([es.Number(et_shp.get</u>('lon')), es.Number(et_shp.get('lat'))]).buffer(1920).bounds()<br>return feature.set(annual_composite.reduceRegion({
      reducer: 'mean',<br>geometry: feature.geometry(),
       scale: 100,
   )\mathcal{W}// Export the FeatureCollection.
Export.table.toDrive({
   collection: mappedFeatures,
   description:
   fileFormat: 'CSV'
});
```
Define the luminosity aggregation function, which takes the centroid and creates a circle buffer around it with a radius that is half the grid size.

Get the average of the luminosity values within the buffer boundary using the **reduceRegion( )** function. *The aggregated luminosity will be stored as a new column in the multipoint shapefile.*

```
. . .
                                                                                  <sup>2</sup> ntl_mean_luminosity.js
var annual_composite = viirs_annual<br>select('b1')//Average DNB radiance values
à.
var nlVis = {
  ar nivis = i<br>min: 0.0,<br>max: 1,<br>bands: ['b1'],
\mathbf{H}Map.centerObject(pt_shp);<br>Map.addLayer(<u>annual_composite.nlVis</u>,"NTL annual composite");
// Aggregate mean of nightlight intensities for centroid regions of size 256px<br>var mappedFeatures = <u>pt_shp.map</u>(function(feature) {<br>var geometry = <u>ee.Geometry.Point(lee.Number(et_shp.get('lon')), ee.Number(et_shp.get('la</u>
      reducer: 'mean',<br>geometry: feature.geometry(),
      scale: 100,
  \overline{\mathcal{W}}\mathcal{W}// Export the FeatureCollection.
Export.table.toDrive({
    collection: mappedFeatures,
   description:
   fileFormat: 'CSV'
});
```
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Export the attribute table of the shapefile as CSV file into the Google Drive.

Copy the codes from the script *ntl_mean_luminosity.js*. Paste the code into the GEE Code Editor, then click **Save**.

If a repository has not yet been created, GEE will prompt to provide a name for the new repository. Click *Create*.

STEP 27

GEE will then prompt to input the script's filename. A description of the script may be provided.

The script will appear in the Script pane.

STEP 28

Click **Assets**.

Click the *Import to script button* to place the NTL into the script.

STEP 30

Rename the variable name from *image* to *viirs_annual*.

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STEP 31

Click the *Import to script button* to place the shapefile into the script.

Rename the variable name from *table* to *pt_shp*.

Scroll to the bottom of the script and locate the section labeled *"Export the FeatureCollection*.*"*

Indicate a filename beside *description*. Then click *Run*.

Click **Tasks**. Note that the task name is the same as the description provided in the output.

STEP 35

Click **Run** to begin processing the code's output.

Verify all the information, including the filename and file format. Ensure that **Drive** is selected to save the output into the Google Drive. Click **Run**.

A check mark will appear to the right of the task name indicating that the task is completed. It may take some time to process.

Go to Google Drive to check for the output CSV file. Download and save the CSV file to the working folder.

STEP 38

From this point, data from the Philippines will be used to illustrate the succeeding steps.

For years without available annual composite imagery, use the Google Earth Engine (GEE). To create the VIIRS annual composite imagery, use the script: *custom_viirs_annual_composite.js*.

Open the JavaScript *custom_viirs_annual_composite.js* using a text editing software (e.g., Windows Notepad) and copy the code.

STEP 40

Paste the code into the GEE code editor then click *Save*.

Change the filter date range and then click **Save**.

STEP 41

GEE will then prompt to input the script's filename. A description of the script may be provided.

The script will appear in the Script pane.

STEP 42

Go to Assets then click the *Import to script button* to place the shapefile into the script.

Rename the variable name from *table* to *pt_shp*.

Step 44

Locate the section labeled *"Export the FeatureCollection"* at the bottom of the script. Indicate a filename beside description then click *Run*.

STEP 45

Go to **Tasks**. Note that the task name is the same as the description provided in the output. Click *Run* to begin processing the code's output.

Verify all the information including the filename and file format. Ensure that *Drive is selected* to save the output into the Google Drive. Click *Run*.

A check mark will appear to the right of the task name indicating that the task is completed. Note that it may take some time to process this task.

Go to Google Drive to verify the output file. Download and save the CSV file to the working folder.

Binning Luminosity Values and Splitting Dataset

Actual nighttime luminosity values are binned into different levels or classes following the approach implemented in the study by Jean et al. (2016) (footnote 1). Binning is done to facilitate more effective training of CNN models. It is implemented using Gaussian mixture models (GMMs). GMMs assume that the distribution of univariate night light intensities comes from the mixture of k-underlying normal or Gaussian distributions and find the set of normal distributions that best fit the data. Based on these, the probability of each observation belonging to each group is derived.

Nighttime luminosity values are grouped into three classes which were found optimal based on experimentation. These are low class, medium class, and high class.

Splitting of datasets is done by performing random sampling within each luminosity bin to preserve overall class distribution. The result is a balanced split of the dataset.

Use the R script **Binning_and_splitting.R** to bin luminosity values.

First, load the required packages.

Select the CSV file containing the average luminosity values.

STEP 3

Set the CSV file's folder path as the working directory.

10 - # get working directory from csv file----11 wd_path <- dirname(NTL_csv_path) 12 # set working directory 13 setwd(wd_path) $1A$

Load the CSV file as the dataframe – *datapoints*.

```
15 \cdot # load csv file to dataframe----
16 datapoints <- read.csv(NTL_csv_path.stringsAsFactors = F)
```
STEP 5

Check the data using the *head( )* function.

```
18 - # check csv data----
19 # please take note of the name of the column containing the luminosity values
20 head(datapoints)
```

```
> # please take note of the name of the column containing the luminosity values
> head(datapoints)
         system.index b1
                                                     filename geocode id
                                                                               lon
                                                                                       lat
1 000000000000000001b0c 0 CNN_DIMG_PHI_2015_ST_384_3840_000313.jpg 148101000 313 121.0896 18.28221
2 000000000000000001b0d    0 CNN_DIMG_PHI_2015_ST_384_3840_000314.jpg 148101000 314 121.1245 18.28221
4 000000000000000001b0f   0 CNN_DIMG_PHI_2015_ST_384_3840_000316.jpg 148101000 316 121.1942 18.28221
5 000000000000000001b10 0 CNN_DIMG_PHI_2015_ST_384_3840_000317.jpg 148105000 317 121.2290 18.28221
6 000000000000000001b11 0 CNN_DIMG_PHI_2015_ST_384_3840_000318.jpg 148105000 318 121.2638 18.28221
                                                               . qeo1 ["type":"Point","coordinates": [121.08963591257178,18.28220730070282]}
2 {"type":"Point","coordinates":[121.12447932820336,18.28220730070282]}
    {"type":"Point","coordinates": [121.159322743835,18.28220730070282]}
3
4 {"type":"Point","coordinates": [121.19416615946663,18.28220730070282]}
5 {"type":"Point","coordinates": [121.22900511599723,18.282207300702822]}
6 {"type":"Point","coordinates": [121.26384853162884,18.28220730070282]}
```
STEP 6

Using the result of *head( )* function, specify the name of the column values and assign it to variable *ntl_col*.

```
22 \times # based on the result of head(), specify the column name containing the average luminosity----
23 ntl_col = "b1"
24
25 # subset column containing the average luminosity
26 avector \leftarrow datapoints[,ntl_col]
```
The luminosity column name that is used by GEE is based on the name of the raster's band, e.g., b1. Generate a subset of this column containing the average luminosity values and store it in the variable *avector*.

```
> # please take note of the name of the column containing the luminosity values
> head(datapoints)
           system.index b1
                                                                   filename geocode id
                                                                                                  lon
                                                                                                             lat
                              CNN_DIMG_PHI_2015_ST_384_3840_000109.jpg 12801000 109 120.9154 18.49126
1 00000000000000001a40
                           Ø
2 00000000000000001a41
                              CNN_DIMG_PHI_2015_ST_384_3840_000110.jpg 12801000 110 120.9503 18.49126
                           \circCNN_DIMG_PHI_2015_ST_384_3840_000135.jpg 12801000 135 120.9154 18.45642
3 00000000000000001a5a 0
4 00000000000000001a5b
                           0 CNN_DIMG_PHI_2015_ST_384_3840_000136.jpg 12801000 136 120.9503 18.45642
5 00000000000000001a75
                           0 CNN_DIMG_PHI_2015_ST_384_3840_000162.jpg 12801000 162 120.8806 18.42158
6 000000000000000001a76 0 CNN_DIMG_PHI_2015_ST_384_3840_000163.jpg 12801000 163 120.9154 18.42158
                                                                               .geo
\begin{array}{ll} \texttt{1} & \texttt{\{``type'':''Point'',''coordinates'': [120.91542329351466,18.49126333539152]\}} \\ \texttt{2} & \texttt{\{``type'':''Point'','' coordinates'': [120.95026670914629,18.49126333539152]\}} \end{array}3 {"type":"Point","coordinates": [120.91542329351466, 18.456419919759895]}
4 {"type":"Point","coordinates": [120.95026670914629, 18.456419919759895]}
5 {"type":"Point","coordinates": [120.88057987788305, 18.42157650412827]}
6
   {"type":"Point","coordinates": [120.91542329351466, 18.42157650412827]}
\geq
```
STEP 7

Use the **class**⁽) function to examine if the extracted luminosity values are of numeric type.

STEP 8

Run the GMM model to produce 3 clusters.

```
30
31 - # run GMM---32 fit=Mclust(avector, G=3, model="V") # request clustering into 3 clusters
33
```
Display the model summary.

```
30
31 - # run GMM---32 fit=Mclust(avector, G=3, model="V") # request clustering into 3 clusters
33
```

```
> # view summary of model ---
> summary(fit)
---------------
                ......................................
Gaussian finite mixture model fitted by EM algorithm
Mclust V (univariate, unequal variance) model with 3 components:
               n df BIC
log-likelihood
                                ICL
      769.2522 35974 8 1454.58 -3717.266
Clustering table:
   1 \quad 2 \quad 325301 8331 2342
×I.
```
Note that there are instances when GMM cannot cluster the data into 2, 3, 4, or 5 clusters because the corresponding cluster distribution is not found. These cases are assumed to be related to country-specific night lights.

Using an if-else statement, determine the course of action that should be taken from the result of the initial GMM calculation.

```
37 * # Check if Mclust yields results ----
38 \cdot if (is.null(fit)=FALSE) {
3940
      # view bins
      fitSclassification
41
42# merge bin results to the original dataframe and select relevant columns
43
44df_bin <- data.frame(datapoints, bin_GMM = fitSclassification) %>%
45
        select(id,
                                            #grid ID
               lon, lat,
46
                                            #centroid coordinates
47
               geocode,
                                            #geocode
48
               avg_rad = all_of(ntl_col), #luminosity values, change column name based on the input csy
               bin_GMM,
49
                                            #bins
50
               filename)
                                            #jpeg filenames
51
52 \cdot }else{ # if the resulting mclust is null
53
54
      #filter luminosity less than or equal to zero
      non_zero_datapoints <- datapoints %>%
55
56
        filter(get(ntl_col)>0)
57
58
      non_zero_avector <- non_zero_datapoints[,ntl_col]
59
60 -# run GMM----
61
      fit=Mclust(non_zero_avector, G=3, model="V") # request clustering into 3 clusters
62
63
      # view summary of model ---
      print(summary(fit))
64
65
      # merge the non-zero luminosity data with its bin classification
66
67
      df_non_zero <- data.frame(non_zero_datapoints, bin_GMM = fitSclassification) %>%
        select(id, bin_GMM) #retain only the id and bin column for ease of merging
68
69
70
      # merge the binned non-zero luminosity data with the rest of the data
71df_bin <- left_join(datapoints,df_non_zero, by="id") %>%
72mutate(bin_GMM = ifelse(is.na(bin_GMM),1,bin_GMM)) %% #classify zero luminosity values into bin category 1
73#grid ID
        select(id,
74lon, lat,
                                            #centroid coordinates
                                            #geocode
75
               geocode,
76
               avg\_rad = all_of(ntl\_col),#luminosity values, change column name based on the input csv
               bin_GMM,
77
                                            #bins
                                            #jpeg filenames
78
               filename)
79 - 1
```
Display the bin classification to check if the initial calculation produced results.

STEP 11

Merge the cluster results with the original dataset. Then select the following relevant columns:

- **id** grid ID,
- **lon, lat** centroid coordinates,
- **geocode** administrative boundary code,
- **avg_rad** luminosity column (renamed to avg_rad),
- **bin_GMM** bin column, and
- **Filename** imagery filename.

```
37 + # Check if Mclust yields results ----
38 \cdot if (is.null(fit)=FALSE)39
40
      # View bins
41fitSclassification
42
43
      # merge bin results to the original dataframe and select relevant columns
      df_bin <- data.frame(datapoints, bin_GMM = fit$classification) %>%
44
45
        select(id,
                                             #grid ID
46
               lon, lat,
                                             #centroid coordinates
47
               geocode,
                                             #geocode
               avg\_rad = all_of(ntl\_col),#luminosity values, change column name based on the input csv
48
               bin_GMM,
49
                                             #bins
50
               filename)
                                             #jpeg filenames
51
```
If the initial calculation yields a null result, generate a subset of the original dataset to extract all positive non-zero luminosity values.

```
52 - \text{lelesf } # \text{ if the resulting mclust is null}53
54
      #filter luminosity less than or equal to zero
55
      non_zero_datapoints <- datapoints %>%
56
        filter(get(ntl_col)>0)
57
      non_zero_avector <- non_zero_datapoints[,ntl_col]
58
59
60 -# run GMM----
      fit=Mclust(non_zero_avector, G=3, model="V") # request clustering into 3 clusters
61
62
63
      # view summary of model---
64
      print(summary(fit))
65
      # merge the non-zero luminosity data with its bin classification
66
67
      df_non_zero <- data.frame(non_zero_datapoints, bin_GMM = fit$classification) %>%
68
        select(id, bin_GMM) #retain only the id and bin column for ease of merging
69
70
      # merge the binned non-zero luminosity data with the rest of the data
      df_bin <- left_join(datapoints,df_non_zero, by="id") %>%
7172
        mutate(bin_GMM = ifelse(is.na(bin_GMM),1,bin_GMM)) %>% #classify zero luminosity values into bin category 1
                                             #grid ID
73
        select(id,
74lon, lat,
                                             #centroid coordinates
75
               geocode,
                                             #geocode
76
               avg\_rad = all_of(ntl\_col),#luminosity values, change column name based on the input csv
               bin_GMM,
77
                                             #bins
78
               filename)
                                             #jpeg filenames
79 - 3
```
Generate another subset of the column containing the average luminosity values and store it in the variable *non_zero_avector*.

```
52 * }else{ # if the resulting mclust is null
53
54
      #filter luminosity less than or equal to zero
55
      non_zero_datapoints <- datapoints %>%
56
        filter(get(ntl_col)>0)
57
58
      non_zero_avector <- non_zero_datapoints[,ntl_col]
59
60 -# run GMM----
      fit-Mclust(non_zero_avector, G=3, model="V") # request clustering into 3 clusters
61
62
63
      # view summary of model---
64
      print(summary(fit))
65
66
      # merge the non-zero luminosity data with its bin classification
67
      df_non_zero <- data.frame(non_zero_datapoints, bin_GMM = fit$classification) %>%
68
        select(id, bin_GMM) #retain only the id and bin column for ease of merging
69
      # merge the binned non-zero luminosity data with the rest of the data
70
71df_bin <- left_join(datapoints, df_non_zero, by="id") %>%
        mutate(bin_GMM = ifelse(is.na(bin_GMM),1,bin_GMM)) %% #classify zero luminosity values into bin category 1
72
73
                                             #grid ID
        select(id,
74
               lon, lat,
                                             #centroid coordinates
               geocode,
75
                                             #geocode
76
               avg\_rad = all_of(ntl\_col),#luminosity values, change column name based on the input csv
77bin_GMM,
                                             #bins
78
               filename)
                                             #jpeg filenames
79 - 1
```
STEP 14

Re-run the GMM model to determine the 3 clusters.

```
52 - \text{leles} # if the resulting mclust is null
53
54
      #filter luminosity less than or equal to zero
55
      non_zero_datapoints <- datapoints %>%
56
        filter(aet(ntl col)>0)
57
58
      non_zero_avector <- non_zero_datapoints[,ntl_col]
59
      # run GMM----
60 -fit=Mclust(non_zero_avector, G=3, model="V") # request clustering into 3 clusters
61
62
63
      # view summary of model---
64
      print(summary(fit))
65
      # merge the non-zero luminosity data with its bin classification
66
67
      df_non_zero <- data.frame(non_zero_datapoints, bin_GMM = fit$classification) %>%
68
        select(id, bin_GMM) #retain only the id and bin column for ease of merging
69
70
      # merge the binned non-zero luminosity data with the rest of the data
      df_bin <- left_join(datapoints,df_non_zero, by="id") %>%
7172mutate(bin_GMM = ifelse(is.na(bin_GMM),1,bin_GMM)) %% #classify zero luminosity values into bin category 1
73
        select(id,
                                             #grid ID
74
               lon, lat,
                                             #centroid coordinates
75
               geocode.
                                             #geocode
76
                avg\_rad = all_of(ntl_col),#luminosity values, change column name based on the input csv
77bin_GMM,
                                             #bins
78
               filename)
                                             #ipeg filenames
79 - 1
```
Print the summary of resulting clusters.

```
52 - }else{ # if the resulting mclust is null
53
54
      #filter luminosity less than or equal to zero
     non_zero_datapoints <- datapoints %>%
55
56
       filter(get(ntl_col)>0)
57
58
      non_zero_avector <- non_zero_datapoints[,ntl_col]
59
60 -# run GMM----
      fit-Mclust(non_zero_avector, G=3, model="V") # request clustering into 3 clusters
61
62
      # view summary of model---
63
64
      print(summary(fit))
65
66
      # merge the non-zero luminosity data with its bin classification
67
      df_non_zero <- data.frame(non_zero_datapoints, bin_GMM = fit$classification) %>%
68
       select(id, bin_GMM) #retain only the id and bin column for ease of merging
69
      # merge the binned non-zero luminosity data with the rest of the data
70
71\,df_bin <- left_join(datapoints,df_non_zero, by="id") %>%
72^{\circ}mutate(bin_GMM = ifelse(is.na(bin_GMM),1,bin_GMM)) %% #classify zero luminosity values into bin category 1
73
        select(id,
                                           #grid ID
74lon, lat,
                                           #centroid coordinates
75geocode,
                                           #geocode
               avg_rad = all_of(ntl_col), #luminosity values, change column name based on the input csv
76
77\,bin_GMM,
                                            #bins
78
               filename)
                                            #ipeg filenames
79 - 3
```

```
Gaussian finite mixture model fitted by EM algorithm
Mclust V (univariate, unequal variance) model with 3 components:
log-likelihood n df
                        BIC
                                  ICL
     -2867.409 2146 8 -5796.19 -6198.36Clustering table:
  \mathbf{1}\overline{2}\overline{\mathbf{3}}1341 581 224
×.
```
Merge the resulting clusters with the non-zero subset and retain only the id and bin_GMM columns.

STEP 17

Merge the binned non-zero dataset with the original dataset using the *left_join( )* function.

```
52 * }else{ # if the resulting mclust is null
53
      #filter luminosity less than or equal to zero
54
55
      non_zero_datapoints <- datapoints %>%
56
        filter(aet(ntl col)>0)
57
58
      non_zero_avector <- non_zero_datapoints[,ntl_col]
59
      # run GMM----
60 -fit=Mclust(non_zero_avector, G=3, model="V") # request clustering into 3 clusters
61
62
      # view summary of model---
63
64
      print(summary(fit))
65
      # merge the non-zero luminosity data with its bin classification
66
67
      df_non_zero <- data.frame(non_zero_datapoints, bin_GMM = fit$classification) %>%
        select(id, bin_GMM) #retain only the id and bin column for ease of merging
68
69
70
      # merge the binned non-zero luminosity data with the rest of the data
      df_bin <- left_join(datapoints,df_non_zero, by="id") %>%
7172
        mutate(bin_GMM = ifelse(is.na(bin_GMM),1,bin_GMM)) %>% #classify zero luminosity values into bin category 1
73
                                             #grid ID
        select(id.
74lon, lat,
                                             #centroid coordinates
75
               geocode.
                                             #geocode
               avg\_rad = all_of(ntl_col),76
                                             #luminosity values, change column name based on the input csv
77bin_GMM,
                                             #bins
78
               filename)
                                             #ipeg filenames
79 - 1
```
Classify all zero luminosity values in cluster 1.

```
52 - }else{ # if the resulting mclust is null
53
54
      #filter luminosity less than or equal to zero
55
      non_zero_datapoints <- datapoints %>%
        filter(get(ntl_col)>0)
56
57
58
      non_zero_avector <- non_zero_datapoints[,ntl_col]
59
60 -# run GMM----
      fit=Mclust(non_zero_avector, G=3, model="V") # request clustering into 3 clusters
61
62
63
      # view summary of model---
64
      print(summary(fit))
65
66
      # merge the non-zero luminosity data with its bin classification
67
      df_non_zero <- data.frame(non_zero_datapoints, bin_GMM = fit$classification) %>%
68
        select(id, bin_GMM) #retain only the id and bin column for ease of merging
69
      # merge the binned non-zero luminosity data with the rest of the data
70
71df_bin <- left_join(datapoints,df_non_zero, by="id") %>%
        mutate(bin_GMM = ifelse(is.na(bin_GMM),1,bin_GMM)) %% #classify zero luminosity values into bin category 1
72^{\circ}73
        select(id,
                                            #grid ID
74lon, lat,
                                            #centroid coordinates
75\,geocode,
                                            #geocode
76
               avg\_rad = all_of(ntl_col),#luminosity values, change column name based on the input csv
77
               bin_GMM,
                                            #bins
78
               filename)
                                            #jpeg filenames
79 - 1
```
STEP 19

Select the relevant columns.

```
52 - }else{ # if the resulting mclust is null
53
54
      #filter luminosity less than or equal to zero
55
      non_zero_datapoints <- datapoints %>%
56
       filter(get(ntl_col)>0)
57
58
      non_zero_avector <- non_zero_datapoints[,ntl_col]
59
60 -# run GMM----
      fit=Mclust(non_zero_avector, G=3, model="V") # request clustering into 3 clusters
61
62
63
      # view summary of model---
64
      print(summary(fit))
65
66
      # merge the non-zero luminosity data with its bin classification
67
      df_non_zero <- data.frame(non_zero_datapoints, bin_GMM = fit$classification) %>%
68
        select(id, bin_GMM) #retain only the id and bin column for ease of merging
69
70
      # merge the binned non-zero luminosity data with the rest of the data
71\,df_bin <- left_join(datapoints, df_non_zero, by="id") %>%
72mutate(bin_GMM = ifelse(is.na(bin_GMM),1,bin_GMM)) %>% #classify zero luminosity values into bin category 1
73
        select(id,
                                             #grid ID
74lon, lat,
                                             #centroid coordinates
75
               geocode.
                                             #geocode
76
               avg\_rad = all_of(ntl_col),#luminosity values, change column name based on the input csv
77
               bin_GMM,
                                             # \hbox{bins}78
               filename)
                                             #ipeg filenames
79
```
Determine the cutoff values for each bin.

```
81 - # Determine the cutoff values for the each bins----
82
    df_{cutoff} < df_{bin} %>%
83
      group_by(bin_GMM) %>%
      summarize(min-cutoff = min(avg\_rad),84
                max_cutoff = max(avg_rad),
85
86
                n\_samples = n()87
88
    # view cutoff table
89 view(df_cutoff)
```


Alternatively, one can use heuristic methods if the GMMs do not provide optimal clusters.

STEP 21

Merge the government-published poverty and population data with the dataset in preparation for machine learning.

Select the government-published dataset.

```
91 - # Merge published poverty and population data-
 92
 93 # select csv file containing published population and poverty data
 94 SAE_csv_path <- tcltk::tk_choose.files(filters = matrix(c("CSV",".csv","All files","*"),2,2,byrow = T),
 95
                                            caption = "Select Published Population and Poverty CSV")
 96 #load csv file as dataframe
 97 df_sae <- read.csv(SAE_csv_path)
 98
 99 # check csv data
100 head(df_sae)
101
102 # merge the dataframe containing binned NTL and published poverty data
103 df \leftarrow left\_join(df\_bin, df\_sae, by = c('geocode'-'PSGC\_code'))104
105 # view merged dataframe
106 head(df)987
```


Load the CSV file as a dataframe.

Assess the structure of the datasets and identify the common variable for joining the two datasets.

```
91 - # Merge published poverty and population data----
 92
93 # select csv file containing published population and poverty data
 94 SAE_csv_path <- tcltk::tk_choose.files(filters = matrix(c("CSV",".csv","All files","*"),2,2,byrow = T),
                                            caption = "Select Published Population and Poverty CSV")
95
 96 #load csv file as dataframe
 97 df_sae <- read.csv(SAE_csv_path)
 98
99
    # check csy data
100 head(df_sae)
101
102 # merge the dataframe containing binned NTL and published poverty data
103 df \leftarrow left\_join(df\_bin, df\_sae, by = c('geocode'-'PSGC\_code'))104
105 # view merged dataframe
106 head(df)107
```


STEP 24

Merge the binned luminosity and government-published datasets using the **left_join( )** function with the geocode and PSGC_code as the join variable.

```
91 - # Merge published poverty and population data--
 92
 93 # select csv file containing published population and poverty data
 94 SAE_csv_path <- tcltk::tk_choose.files(filters = matrix(c("CSV",".csv","All files","*"),2,2,byrow = T),
                                           caption = "Select Published Population and Poverty CSV")
 95
 96 #load csv file as dataframe
 97 df_sae <- read.csv(SAE_csv_path)
 98
99 # check csv data
100 head(df_sae)
101
102 # merge the dataframe containing binned NTL and published poverty data
103 df <- left_join(df_bin,df_sae, by = c('geocode'='PSGC_code'))
104
105 # view merged dataframe
   head(df)
106
107
```


Check the structure of the new dataset structure to ensure that the two datasets are merged.

```
91 - # Merge published poverty and population data----
9293 # select csv file containing published population and poverty data
94 SAE_csv_path <- tcltk::tk_choose.files(filters = matrix(c("CSV",".csv","All files","*"),2,2,byrow = T),
                                           caption = "Select Published Population and Poverty CSV")
95
96 #load csv file as dataframe
97 df_sae <- read.csv(SAE_csv_path)
9899 # check csv data
100 head(df_sae)
101
102 # merge the dataframe containing binned NTL and published poverty data
103 df \leftarrow left\_join(df\_bin, df\_sae, by = c('geocode'-'PSGC\_code'))104
105 # view merged dataframe
106 head(df)
107
```


Split the dataset into training and test sets. It is up to the user to decide on an optimal splitting strategy. In the ADB study (footnote 2), the dataset was split into two: 90% for training and 10% for test. The training dataset will be used for training the CNN model. This dataset is further split into 80% for training and 20% for validation through fastai. After developing the trained model, the test dataset will be used to validate its accuracy.

First, load the package caret. This package contains the function **createDataPartition( )** that will enable the generation of a balanced split in the dataset. **createDataPartition( )** returns the row index of the dataset belonging to the specified split.

```
108 - # Dataset Splitting-----
109 # Data shall be split into 90% for training and validation and 10% holdout dataset
110 library(caret)
111
112 #generate index of the 90% training and validation dataset
113 splitIndex <- createDataPartition(df$bin_GMM,
                                                         #specify column for basis of split
                                                         #number of split
114
                                        times = 1,
115
                                        p = 0.9#percent split
116
                                        list=FALSE) #outputs the data as a matrix
117
118 #subset dataset to extract the training and validation dataset
119 df_{\text{train}} \leftarrow df_{\text{splitIndex}}120 #subset dataset to extract the holdout dataset
121 df_{\text{Test}} \leftarrow df[\text{-splitIndex},]122
123 #check the resulting datasets
124 head(df_Train)
125 head(df_Test)
126
127 nrow(df_Train)
128 nrow(df_Test)
```
createDataPartition( ) requires the following parameters:

- column of dataset for the basis of the split,
- *times* number of split to perform, in our case only one,
- *p* − split ratio in our case 0.9 or 90%, and
- **list = FALSE** to output the data as a matrix. This will be used when subsetting the dataset.

```
108 - # Dataset Splitting-----
109 # Data shall be split into 90% for training and validation and 10% holdout dataset
110 library(caret)
111
112 #generate index of the 90% training and validation dataset
113 splitIndex <- createDataPartition(dfSbin_GMM,
                                                   #specify column for basis of split
114
                                      times = 1,#number of split
115
                                      p = 0.9,
                                                       #percent split
116
                                                      #outputs the data as a matrix
                                      list = FALSE)117
118 #subset dataset to extract the training and validation dataset
119 df_Train <- df[ splitIndex,]
120 #subset dataset to extract the holdout dataset
121 df_Test <- df[-splitIndex,]
122
123 #check the resulting datasets
124 head(df Train)
125 head(df_Test)
126
127 nrow(df_Train)
128 nrow(df_Test)
```
STEP 28

Extract the training and test datasets from the subset of the dataset.

```
108 - # Dataset Splitting-----
109 # Data shall be split into 90% for training and validation and 10% holdout dataset
110 library(caret)
111
112 #generate index of the 90% training and validation dataset
113 splitIndex <- createDataPartition(dfSbin_GMM, #specify column for basis of split
                                                        #number of split
114
                                       times = 1,115
                                       p = 0.9,
                                                     #percent split
116
                                       list=FALSE) #outputs the data as a matrix
117
118 #subset dataset to extract the training and validation dataset
119 df_Train <- df[ splitIndex,]
120 #subset dataset to extract the holdout dataset
121 df_{\text{Test}} \leftarrow df_{\text{splitIndex}}122123 #check the resulting datasets
124 head(df_Train)
125 head(df_Test)
126
127 nrow(df_Train)
128 nrow(df_Test)
```
Check the dataset's structure.

 $\frac{1}{2}$ #check the resulting datasets

```
108 - # Dataset Splitting-----
109 # Data shall be split into 90% for training and validation and 10% holdout dataset
110 library(caret)
111
112 #generate index of the 90% training and validation dataset
113 splitIndex <- createDataPartition(dfSbin_GMM,
                                                        #specify column for basis of split
114
                                         times = 1,#number of split
115
                                         p = 0.9,
                                                           #percent split
116
                                         list = FALSE)#outputs the data as a matrix
117
118 #subset dataset to extract the training and validation dataset
119 df_{\text{train}} < df_{\text{split}} splitIndex,
120 #subset dataset to extract the holdout dataset
121 df_{\text{Test}} \leftarrow df_{\text{self}} \leftarrow ff_{\text{self}}122
123 #check the resulting datasets
124 head(df_Train)
125 head(df_Test)
126
127 nrow(df_Train)
128 nrow(df_Test)
```


Check the number of observations per dataset by displaying the number of rows.

```
108 - # Dataset Splitting-----
109 # Data shall be split into 90% for training and validation and 10% holdout dataset
110 library(caret)
111
112 #generate index of the 90% training and validation dataset
113 splitIndex <- createDataPartition(dfSbin_GMM, #specify column for basis of split
                                     times = 1,#number of split
114
115
                                     p = 0.9,
                                                   #percent split
116
                                     list=FALSE) #outputs the data as a matrix
117
118 #subset dataset to extract the training and validation dataset
119 df_Train <- df[ splitIndex,]
120 #subset dataset to extract the holdout dataset
121 df_Test <- df[-splitIndex,]
122
123 #check the resulting datasets
124 head(df_Train)
125 head(df_Test)
126
127 nrow(df_Train)
128 nrow(df_Test)
```

```
> nrow(df_Train)
\lceil 1 \rceil 18081
> nrow(df_Test)
[1] 2009
\geq
```
STEP 31

Output the two datasets as CSV files.

```
130 \cdot # output results to as csv files----
131 # generate filename
132 train_file_name <- str_replace(basename(NTL_csv_path), "full", "train90")
133 test_file_name <- str_replace(basename(NTL_csv_path), "full","test10")
134
135 write.csv(df_Train, train_file_name, row.names = F)
136 write.csv(df_Test, test_file_name, row.names = F)
```
Upload the files in Google Drive. This will be used for training the CNN model.

4 TRAINING OF CONVOLUTIONAL NEURAL NETWORK

Aconvolutional neural network (CNN) is a subclass of artificial neural networks that is primarily used
in computer vision (e.g., classification, recognition). It is designed to cope with a large amount of unstructured and pixelated data from digital images. In this context, a CNN is trained to extract features in daytime images using intensity of night lights as labels. These extracted features are then used to predict poverty.

STEP 1

In the browser address bar, input the Google Colab (footnote 7) web address [https://colab.research.](https://colab.research.google.com/) [google.com/](https://colab.research.google.com/) and press **Enter** from the keyboard. *Make sure to log in to Google account*. Then click **Upload**.

Click **Choose File**.

Locate the Jupyter Notebook file from the computer. Use **CNN_training_template.ipynb**. Click **Open**.

Setup the runtime type once the file has loaded. Click **Runtime** on the menu bar.

Then click **Change runtime type**.

On the Notebook settings, change **Hardware accelerator** into **GPU**. Then click **Save**.

STEP 5

Click **Connect**.

This will initialize the Colab's environment.

STEP 6

To execute, click each code cell and click \bullet button at the beginning of each cell.

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Setup and mount the Google Drive (footnote 6).

STEP 7

In the browser, sign in to your Google account.

Click **Allow**.

Click the **Copy** icon \Box to copy the code.

Return to the Colab browser tab. **Paste** the code in the text box. Then press **Enter**.

A status will show the path where Google Drive is mounted.

STEP 9

Ensure that modules are reloaded automatically and any charts or images displayed are shown in this notebook.

STEP 10

Locate the path to the CSV file containing the binned luminosity values that was previously uploaded in Google Drive.

Click **Files** icon **D** to show the **Files section**.

STEP 12

Click **gdrive** from the list of folders and expand the file directory tree to find the CSV file location.

Click the vertical ellipsis to show more file options.

STEP 14

Click **Copy path**.

Paste the link on the blank space after the variable *csv_path* and enclose in apostrophes.

STEP 16

Execute the code cell to check the contents of the first five rows of the CSV file.

The information on the column contents will be used later in building the ImageDataBunch object, particularly the binned luminosity and filename column.

Import **os** and **shutil** python modules and create folder **data** in the Colab virtual machine's drive.

STEP 18

Click **Files** icon **the Solut** to show the **Files section**.

From the list of folders, click *gdrive*.

Expand the file directory tree to find the location of the *tar.gz* file.

Click the vertical ellipsis to show more file options.

STEP 21

Click **Copy path**.

Paste the link beside the variable *tar_file* and enclose it in apostrophes.

STEP 23

Count the number of daytime imagery files extracted.

STEP 24

The CNN training process starts in this step.

Import all the necessary packages in fastai.

STEP 25

Check the fastai version to determine if the latest version is running.

STEP 26

Define all the parameter variables needed to create the ImageDataBunch. Load **re** library to be used for string manipulation.

The *root_col* variable stores the root directory path containing the daytime satellite images. The *valid_pct* command stores the percentage of dataset used for validation.

From the previous code, check the data contained in the CSV file, particularly the *bin_GMM* and *filename*.

The *label_col* command stores the name of binned-luminosity-containing column. The *filename_col* command stores the name of the filename-containing column.

STEP 29

Extract the country code, year, daytime satellite imagery code, and imagery file resolution from the *tar.gz* filename. Then store them in variables *country, year, day_sat*, and *img_res*, respectively.

Generate and print the filename to be used when saving the learner and model objects.

STEP 31

Define the image transformation to be applied to the daytime images, like vertical flipping, random lighting and contrast change with 10% probability, dihedral and symmetric warp. *This is called data augmentation. Data augmentation is used to increase the number of samples in the training dataset, to get the model to generalize better, and to mitigate imbalanced classes in dataset. It also prevents the model from overfitting. In effect, it increases the accuracy of the model.*

Define the ImageDataBunch.

ImageDataBunch is a fastai object, which stores the path to the image folder, training dataset, augmentation, and other settings of the training.

STEP 33

View the first 25 images of the training dataset.

Create a CNN learner object with the pre-trained model, training and validation datasets, metrics, and loss function as arguments. A **model** is the combination of mathematical functions and parameters or weights. Both **metrics** and **loss** functions measure the model's performance, but they differ in use. **Metrics** are used by researchers to define the performance of their models, while **loss** functions are used by the deep learning platform to update the model's weights during training.⁹

Set the CNN model parameter to ResNet-34 and metrics to **error_rate**. Resnet models have been trained on an image-net database of over 14 million images, with 1.2 million of them assigned to one of a thousand categories. It has different variants like ResNet-18, ResNet-34, ResNet-50, ResNet-101, ResNet-110, and ResNet-152, which differ in the number of layers. According to PyTorch documentation (<https://pytorch.org/docs/stable/torchvision/models.html>), ResNet-34 has higher accuracy and six times fewer parameters compared to the pre-trained model VGG. The reduced file size of ResNet-34 is important since no dedicated stand-alone hardware is used for training the model. Though ResNet-18 has smaller number of parameters and smaller file size, ResNet-34 performs better.

The learner also uses a **weighted Cross Entropy loss** function to mitigate imbalanced prediction classes. It penalizes the model for wrong prediction of low frequency class (i.e., 3- high nightlight) based on weight. It also prevents the model from tending to predict more of low nightlight classes 1 and 2 because these classes have the most samples. Weights [0.7,1.0,1.1] are chosen based on experiments. In general, however, users may define other weights as deem suitable (see Box 1).

STEP 35

Define the callbacks. In fastai, **callbacks** are functions that are executed when an "event" occurs during the training process.

⁹ "Lesson 2 - Deep Learning for Coders (2020)", Youtube video, 1:31:04, posted by Jeremy Howard on 22 August 2020. https://www.youtube.com/watch?v=BvHmRx14HQ8.

The first callback function saves the weights of the best training cycle in the batch into a **.pth** file with specified filename.

The second callback function displays a graph of training and validation dataset loss during training.

The last callback function stops the training batch when there are three consecutive training cycles that did not improve the model.

STEP 36

Execute the code to train the model using the dataset. Since the pre-trained CNN is used, the weights are already in place and thus the number of training epochs can be lower. An **epoch** is equal to one cycle of training through all the training dataset.

Unfreeze the last layer group and train it for 14 epochs. The layer group being trained will determine the final predictions. This will create new weights for the layer group that will identify what an image looks like if it belongs to either of the three luminosity intensity classes (i.e., 1=low, 2=medium, 3=high).

A higher epoch can be used, however, a point will be reached when the errors no longer change. Even if the training continues further, the last best model will still be saved through the first callback function. Also, as specified in the third callback function, the training stops after three consecutive cycles without the model improving. This will save time and computing resources.

A weight decay of 0.1 is also used, following the best practice for fastai as suggested by its developers. **Weight decay** is a model regularization technique where it penalizes parameters (weights) to prevent overfitting. Too large a weight decay could prevent the model from fitting well, in other words, the model is not "learning". Too small a weight will make the model over-fit earlier.¹⁰

Upon execution, the following will be displayed:

 \blacksquare tabulated training, validation loss, and error rate per training cycle (epoch),

¹⁰ "Lesson 5: Deep Learning 2019 - Back propagation; Accelerated SGD; Neural net from scratch", Youtube video, 2:13:33, posted by Jeremy Howard on 26 January 2019. [https://www.youtube.com/watch?v=CJKnDu2dxOE](https://www.youtube.com/watch%3Fv%3DCJKnDu2dxOE).

training and validation loss graph, which is the second callback function, and

■ resulting models with better error_rate from each epoch.

Unfreeze the last two layer groups of the model.

Find the best learning rate. The **learning rate** specifies the degree of change of the parameters. The parameters are adjusted based on the gradient to decrease the loss function. A **cyclical learning rate** approach eliminates the need to experimentally find the best values and schedule for the global learning rates. Instead of monotonously decreasing the learning rate, this method lets the learning rate cyclically vary between reasonable boundary values. Training with cyclical learning rates instead of fixed values achieves improved classification accuracy without the need to fine-tune and iterate.

Plot the best learning rate.

Take note of the range of learning rate before the loss starts to rise.

Unfreeze the last two layer groups.

Train for six more epochs.

 \bullet learn.freeze_to(-2) learn.fit_one_cycle(6, max_lr=slice(1e-6, 1e-3), wd=0.1) D epoch train_loss valid_loss error_rate time $\mathbf{0}$ 0.410242 0.498959 0.153333 00:39 0.418009 0.522044 0.173333 00:39 \mathbf{I} $\overline{2}$ 0.430769 0.475751 0.166667 00:40 0.417897 0.493684 3 0.168333 00:40 0.393658 $\overline{4}$ 0.455786 0.153333 00:40 5 0.378205 0.455251 0.153333 00:40 Better model found at epoch 0 with error_rate value: 0.15333333611488342. train valid 0.5 0.4 -0.3 0.2 0.1 0.0 $\overline{25}$ 50 75 100 125 150 175 200

Specify the learning rate range generated from the previous graph.

STEP 39

Define the interpretation methods for classification models. Generate a confusion matrix and visualization of the images with inconsistencies. *A confusion matrix or error matrix can validate and enhance the performance of the machine learning classification-related tasks by comparing the number of correct and incorrect predicted images and employing a particular loss function to minimize imbalanced prediction losses.*

Extract the top losses and the corresponding image ID.

Check if the validation dataset, losses, and image IDs (idx) are of the same number.

STEP 40

Plot the satellite images with highest training losses or with inconsistencies.

Take note of any inconsistences between the input data and the output class (e.g., low-quality day images, high percentage of cloud cover, or illogical nightlight category).

Print the corresponding image filenames of satellite images with high loss function values. In this example, the filenames of the top 50 satellite images with high loss function values are displayed.

Plot the confusion matrix to further validate the training process. On the vertical axis, list the known classes for each image, in this case the nighttime light intensity. On the horizontal axis, list the predictions from the CNN. Each cell contains the number of images for true and predictive classes. Correctly predicted images lie on the main diagonal and every other image lies on the off diagonal. As the classes are ordinal (class1 < class2 < class3: low < middle < high intensity), it holds that the farther away the values are from the main diagonal, the larger the error. (Note: Other projects might have non-ordered classes like "cats versus dogs", hence, the distance to the diagonal is irrelevant.) These values should be as small as possible to avoid "big mistakes" during prediction.

Present the list of largest non-diagonal entries of the confusion matrix. *This refers to actual, predicted, and number of occurrences.*

Box 1. Steps in Adjusting Weights of Cross Entropy Loss Function

- 1. Start with equal weights of [1.0, 1.0, 1.0].
- 2. Unfreeze the last layer and train for 14 epochs.
- 3. Plot and check the confusion matrix results.

Try to achieve a relatively balanced matrix.

- In Figure A, the equal weights created a confusion matrix with more predictions below the diagonal.
- In Figure B, the extreme low and extreme high 1st and 3rd weights are tried, respectively. This resulted in a higher prediction above the diagonal.
- In Figure C, a relatively balanced matrix is achieved.

Define the function for removing "anomalous" images from the training and validation dataframe.

If there is a significant number of inconsistencies between input data and output class (e.g., low-quality daytime images, too cloudy images), remove these instances from the original dataframe. *Since the ImageDataBunch contains labels and image file path, remove these images using their filenames as subset parameters for the dataframe.*

```
#Function for dropping images from dataframe
def drop image(loss index):
  filename list = [os.path.basename(data.valid ds.x.items[i]) for i in loss index]
  # view data to be dropped
  print(df.loc[df['filename'].isin(filename list)])
  # get filename and index of rows to be dropped from dataframe
  df filenames = df['filename'].loc[df['filename'].isin(filename list)]index_names = df.loc[df['filename'].isin(filename_list)].index
  df.drop/index\_names, in place = True)print("Image filenames dropped from dataframe:")
  for f in df filenames:
    print(f)
```
STEP 44

Print the indexes of the images belonging to the top 50 highest losses. Based on the image plot of the 50 top losses, select the "anomalous" images to be removed. *Note that this step is optional.*

```
print("Row index of top 50 losses:")
      print(idxs)\Gamma Row index of top 50 losses:
     tensor([1165, 2050, 288, 1032, 2226, 871, 2365, 1227, 1020, 2252, 21, 38,<br>1374, 2367, 1461, 229, 603, 1581, 1868, 2157, 926, 1453, 1959, 2071,<br>11, 1061, 1256, 1177, 492, 2371, 2211, 1822, 424, 1837, 244, 907,
                    320, 2145, 481, 1485, 1170, 2161, 1810, 2146, 98,
                                                                                                         20, 628, 2063,
                  1955, 1343])
```


Assign the selection as a list data type to the variable *selected_index*. Call the *drop_image( )* function to pass the index of images to be dropped.

```
[28] print("Row index of top 50 losses:")
     print(idxs)\Gamma Row index of top 50 losses:
     tensor([1165, 2050, 288, 1032, 2226, 871, 2365, 1227, 1020, 2252, 21,
                                                                                          38,1374, 2367, 1461, 229, 603, 1581, 1868, 2157, 926, 1453, 1959, 2071,
               11, 1061, 1256, 1177, 492, 2371, 2211, 1822, 424, 1837, 244, 907, 320, 2145, 481, 1485, 1170, 2161, 1810, 2146, 98, 20, 628, 2063,
              1955, 1343])
Selected_index=[2050,1032,244,2146,98,20,628,2063]
     drop_image(selected_index)
```
Execute the code cell.

The function will print out the data associated with the images.

Confirm the filenames of the images.

After removing the "anomalous" data, repeat steps to generate a ImageDataBunch, creating learner and training for 14 epochs with the dataset.

STEP 48

Unfreeze the last three layer groups of the model. Find the best learning rate and plot it.

Unfreeze the last three layer groups and train for six more epochs using the learning rate range determined from the previous graph.

In this scenario, note that the model did not improve after three cycles, thus the training was terminated.

Unfreeze all layer groups and determine the best learning rate again.

STEP 51

Unfreeze all the layers and train for three more epochs using the learning rate from the previous graph. *This step ensures the consistency of the whole network.*

Define again the interpretation methods for classification of models. Extract the top losses and the corresponding image ID. Lastly, check if the validation dataset, losses, and image IDs (idx) are of the same length.

STEP 53

View the images again showing the top losses from the model's prediction, actual value, training loss, and probability.

Generate the confusion matrix to validate the training process.

STEP 55

Save the learner object and model weights in Google Drive.

```
learn.export(file=learner_filename) #train and export learner
\bulletlearn.save(modelWt_filename)
    # define folders
    save_path = "/content/gdrive/MyDrive/models/"
    os.makedirs(save_path, exist_ok=True)
    shutil.copy(os.path.join('/content/',learner_filename), save_path)
    shutil.copy(os.path.join("/content/models/",modelWt_filename+'.pth'), save_path)
```
Test the trained CNN model using the 10% test dataset.

First, clear the virtual memory.

Prepare the ImageDataBunch for the test dataset and load the trained CNN model and learner objects.

```
\bullet bs val = 64
                     # batch size, change to 16 if you run out of memory even after clicking Kernel->Restart
    #create Databunch
    df = pd.read_csv(test_dataset) #load test dataset with holdout images(10%) and labels(classes)
    df_val = df[['bin GMM', 'filename']]#create ImageList with folder of all images and dataset of filenames and corresponding classes of our test set
    img_list = ImageList.from_df(df=df_val, path='/content/data/', cols='filename', folder=imagery_folder, suffix='')
    img_list_split = img_list.split_none() #all data on train set, not splitting to train and validation sets like in databunch
    list_label = img_list_split.label_from_df(0)
    list_label.valid = list_label.train # trick: load training dataset as validation dataset
    print(list_label) #check what is inside train, validation and test set at the moment
    list label.transform(tfms=None.size=int(img_res)) # optional transforms
    data = list_label.databunch(bs=bs_val);
    data.normalize(imagenet stats)
    learn = cnn_learner(data, models.resnet34, metrics = error_rate)
    learn = load_learner('/content/', file=learner_filename) #learner object must be used for inference purposes
    learn.load(modelWt_filename) #load weights of the model, which we want to test
    learn.data.valid dl = data.valid dl #override with inference data with transfroms and other..
    learn.loss func = torch.nn.CrossEntropyLoss()
    learn.metrics #check which metrics is set up
    interp = ClassificationInterpretation.from_learner(learn,ds_type=DatasetType.Valid) #perform interpretation for validation
    interp.plot_confusion_matrix() #matrix representing predictions on holdout test set
```
STEP 58

The code cell will output information regarding the data split and confusion matrix.

Plot the top 25 images with high losses and overlay a heatmap to indicate areas in the images that the CNN considers as important for actual nightlight class.

Then define the *evaluate_model_from_interp( )* function to evaluate the overall accuracy of the model.

```
tfms = None\bulletdata_test = datadef evaluate_model_from_interp(interp, data):
         # perform evaluation of the model to take a look at predictions vs. labels and compute accuracy
         print(f'Interp has {len(interp.y_true)} ground truth labels: {interp.y_true}')
         print(f'Interp yielded {len(interp.preds)} raw predictions. First two raw predictions are: {interp.preds[:2]}')
         print(f'The problem had {len(data.classes)} classes: {data.classes}') # data.c is just len(data.classes)
       print('')print(f'Pred -> GroundTruth = PredLabel -> GroundTruthLabel')
        ok pred = 0for idx, raw_p in enumerate(interp.preds):
              pred = np.argv(ax(xaw_p))if idx < 10: #display first 10 predictions and corresponding real labels
                  print(f'(pred) \rightarrow (interp.y_ttrue_idx]) = (data.classes(pred)) \rightarrow (data.values_i(y_idx));if pred == interp.y_true[idx]: #count correct predictions
                   ok pred += 1acc = ok_pred / len(interp.y_true) #calculate accuracy by correct predictions divided by total predictions
         print(f'Overall accuracy of the model: {acc:0.5f}')
    #call function
    evaluate_model_from_interp(interp, data_test)
\lbrack. \rbrack Interp has 1323 ground truth labels: tensor([0, 0, 0, ..., 0, 0, 0])
    Interp yielded 1323 raw predictions. First two raw predictions are: tensor([[9.6205e-01, 3.7922e-02, 3.2040e-05],
              (9.9941e-01, 5.9113e-04, 3.3564e-08]]The problem had 3 classes: [1, 2, 3]
    Pred \rightarrow GroundTruth = PredLabel \rightarrow GroundTruthLabel\begin{array}{cccc} 0 & \!\!-\!\!> & \!\!0 & \!\! = & \!\!1 & \!\!-\!\!> & \!\!1 \\ 0 & \!\!-\!\!> & \!\!0 & \!\! = & \!\!1 & \!\!-\!\!> & \!\!1 \end{array}0 \Rightarrow 0 = 1 \Rightarrow 10 \Rightarrow 0 = 1 \Rightarrow 1Overall accuracy of the model: 0.93197
```


5 CONVOLUTIONAL
5 NEURAL NETWORK NEURAL NETWORK MODEL FEATURE EXTRACTION

After training the CNN, the next step is to extract the abstract satellite image features that are correlated with the intensity of night lights. This is done by altering the model such that it generates the features from the last hidden layer as an output rather than as a regular classification category output. In this case, the feature vectors that the CNN uses to specify the intensity of night lights are extracted.

Data Requirements

- Archive file containing daytime satellite imagery (JPG)
- CSV file containing binned luminosity values and government-published poverty estimates
- Trained CNN model

Tools

Google Colaboratory (footnote 7) (CNN_training_template.ipynb)

STEP 1

For feature extraction, open a new notebook file. Click **File**.

Then click **Upload Notebook**.

STEP 2

Click **Choose File**. Use the Jupyter Notebook file **CNN_feature_extraction.ipynb**.

Locate the file and Click **Open**.

STEP 3

Setup the runtime type once the file has loaded. Click **Runtime** on the menu bar.

On the Notebook settings, change **Hardware accelerator** into **GPU**. Then click **Save**.

Click **Connect**.

This will initialize the Colab environment.

For environment setup, mount Google Drive (footnote 6) to Google Colab.

STEP 6

In the browser, sign in to Google account.

Click **Allow**.

Click **Copy** icon $\boxed{\Gamma}$ to copy the code.

Return to the Colab browser tab. **Paste** the code in the text box. Then press **Enter**.

A status will show the path where Google Drive is mounted.

STEP 7

Ensure that modules are reloaded automatically and any charts or images displayed are shown in the notebook.

STEP 8

Locate the path of the training dataset's CSV file.

```
#paste into link the link of csv file
import pandas as pd
train dataset = ''test_dataset = train_dataset.replace('train90','test10')
df_train = pd.read_csv(train_dataset)
df_test = pd.read_csv(test_dataset)
```
Click Files icon **the Solution** to show the **Files section**.

STEP 10

Click **gdrive** from the list of folders and expand the file directory tree to find the CSV file location.

Click the vertical ellipsis to show more file options.

Click **Copy path**.

STEP 13

Paste the link on the blank space after the variable *train_dataset* and enclose in apostrophes.

Create an identifying column in the training and test datasets, merge the two, sort the dataframe by grid ID, and print out the first four rows of the dataset.

STEP 15

Load **os** and **shutil** packages for operating system functionality and for unpacking archive files, respectively.

Click **Files** icon **the Solut** to show the **Files section**.

STEP 17

From the list of folders, click *gdrive* and expand the file directory tree to find the targ.gz file location.

Click the vertical ellipsis to show more file options.

Click **Copy path**.

STEP 20

Paste the link beside the variable *tar_file* and enclose it in apostrophes.

Generate the different parameters for the CNN model.

```
O import re
    # Specify how much of the network are we cutting away,
    # NOTE: this does not correspond to single layers but smaller components (Linear weights, RELU and others)
    layer\_drops = 2#extract country code, year, daytime satellite imagery source and imagery file resolution from tar filename
    country, year, day_sat, img_res = re.search("[A-2]{3}_[0-9]{4}_[A-2]{2}_[0-9]{3}",tar_file).group().split("_")
    target variable name = "POV "+ year
```
Check if all satellite imagery in the CSV file are present in the folder.

STEP 22

Delete the rows in the dataframe that do not have a corresponding imagery, otherwise fastai's databunch will not work.

Define the necessary parameters for creating ImageDataBunch.

STEP 24

Import all libraries that are needed for the extraction of features from the trained CNN model.

STEP 25

Load the dataset to the ImageDataBunch.

STEP 26

Create a learner object from the fastai library containing the datasets (i.e., images and labels) without the pre-trained CNN.

Copy the pre-trained model from Google Drive to the Google Colab virtual machine drive.

```
\bullet# define gdrive CNN model save path
    source path = "/content/gdrive/MyDrive/models/"
    shutil.copy(os.path.join(source_path,learner_filename), root_col)
    shutil.copy(os.path.join(source path,modelWt filename+'.pth'), root_col)
```
STEP 28

Load the trained CNN model and merge it with the dataset in the learner object. It also outputs the ImageDataBunch information and structure of the model layers.

```
learn.load(root col + modelWt filename)
[+ Learner(data=ImageDataBunch;
   Train: LabelList (16072 items)
   x: ImageList
   Image (3, 384, 384), Image (3, 384, 384), Image (3, 384, 384), Image (3, 384, 384), Image (3, 384, 384)
    y: CategoryList
   1, 1, 1, 1, 1Path: /content;
   Valid: LabelList (4018 items)
    x: ImageList
    Image (3, 384, 384), Image (3, 384, 384), Image (3, 384, 384), Image (3, 384, 384), Image (3, 384, 384)
    y: CategoryList
    1, 1, 1, 1, 1Path: /content;
    Test: None, model=Sequential(
      (0): Sequential(
        (0): Conv2d(3, 64, kernel_size=(7, 7), stride=(2, 2), padding=(3, 3), bias=False)
        (1): BatchNorm2d(64, eps=1e-05, momentum=0.1, affine=True, track_running_stats=True)
        (2): ReLU(inplace=True)
        (3): MaxPool2d(kernel_size=3, stride=2, padding=1, dilation=1, ceil_mode=False)
        (4): Sequential(
         (0): BasicBlock(
            (conv1): Conv2d(64, 64, Kernel_size=(3, 3), stride=(1, 1), padding=(1, 1), bias=False)(bnl): BatchNorm2d(64, eps=1e-05, momentum=0.1, affine=True, track_running_stats=True)
            (relu): ReLU(inplace=True)
            \text{(conv2): Conv2d(64, 64, Kernel_size=(3, 3), stride=(1, 1), padding=(1, 1), bias=False)}(bn2): BatchNorm2d(64, eps=1e-05, momentum=0.1, affine=True, track_running_stats=True)
          \lambda
```
Select two test images from the dataframe and load them into the python environment. This is helpful when trying out functions that operate on images.

STEP 30

Insert the predict function as a method of the learner class. This method returns only the node values of the last layer in the model, which are normally probabilities of each output category.

STEP 31

Compare the result of the predict function with the custom predict function that was previously defined.

```
print(learn. predict (pic_one))
print(learn.my_predict (pic_one))
(Category tensor(0), tensor(0), tensor([1.0000e+00, 1.6193e-06, 2.8243e-09]))
tensor([1.0000e+00, 1.6193e-06, 2.8243e-09])
```
Generate a new model without the last fully connected layer.

```
Ω
   new model = learn
   print('Original fully-connected layer group length:
                                                              + str(len(learn.model[1])))
   print('-------------')
   print("Original fully-connected layer structure:")
   print(learn.model[1])
   print('')print('')new model.model[1] = new model.model[1][:-layer drops]
   print('New fully-connected layer group length: ' + str(len(new model.model[1])))
   print('-----')print("New fully-connected layer structure:")
   print(new_model.model[1])
[3 Original fully-connected layer group length: 9
   Original fully-connected layer structure:
   Sequential(
      (0): AdaptiveConcatPool2d(
       (ap): AdaptiveAvgPool2d(output size=1)
       (mp): AdaptiveMaxPool2d(output_size=1)
     (1): Flatten()(2): BatchNormld(1024, eps=1e-05, momentum=0.1, affine=True, track_running_stats=True)
     (3): Dropout(p=0.25, inplace=False)
     (4): Linear(in_features=1024, out_features=512, bias=True)
     (5): ReLU(inplace=True)
     (6): BatchNormld(512, eps=1e-05, momentum=0.1, affine=True, track_running_stats=True)
     (7): Dropout(p=0.5, inplace=False)
      (8): Linear(in_features=512, out_features=3, bias=True)
   \overline{1}New fully-connected layer group length: 7
   New fully-connected layer structure:
   Sequential(
     (0): AdaptiveConcatPool2d(
       (ap): AdaptiveAvgPool2d(output_size=1)
       (mp): AdaptiveMaxPool2d(output_size=1)
     \rightarrow(1): Flatten()(2): BatchNormld(1024, eps=1e-05, momentum=0.1, affine=True, track_running_stats=True)
     (3): Dropout(p=0.25, inplace=False)
     (4): Linear(in_features=1024, out_features=512, bias=True)
     (5): ReLU(inplace=True)
     (6): BatchNormld(512, eps=1e-05, momentum=0.1, affine=True, track_running_stats=True)
   \lambda
```
Define a new function that extracts the tensor of the image features. Then measure the tensor length. **Tensors** are multidimensional arrays. It functions like a numpy array however it has an added benefit where it can be calculated on a graphics processing unit.⁹

STEP 34

Before predicting image features, create an empty array for storing extracted features and a dataframe containing image file names.

STEP 35

Loop through the images and extract the features.

```
Œ
   for i, path i in enumerate (features out img):
      # open the image with the fastai open image function
      temp_img = open_image(os.path.join(imagery_path, path_i))
      # extract the features of the single image
      tempfeatures = Extract_Features (temp_img).flatten().reshape(1, -1).numpy()
      # store them for output
      features_out[i,:] = tempfeatures
```
Merge the extracted features with the image file names.

 \bullet features_out_pd = pd.DataFrame(data = features_out, index = features_out_img)

STEP 37

Save the CSV file to Google Drive, which will be used for ridge regression.

```
\bulletcsv_path ="/content/gdrive/MyDrive/"
     features_filename = "_".join(["CNN_FOUT_RES34", country, year, day_sat, str(img_res)])+".csv"<br>CENI_full_filename = "_".join(["CNN_CENI_RES34", country, year, day_sat, str(img_res)])+".csv"
[ ] # save to disk
     features_out_pd.to_csv(features_filename)
     df_full.to_csv(CENI_full_filename)
     # copy from colab virtual drive to google drive
     shutil.copy(os.path.join('/content/',features_filename), csv_path)
     shutil.copy(os.path.join('/content/',CENI_full_filename), csv_path)
```
6 RIDGE REGRESSION

In the final training step, ridge regression is implemented to determine the relationship between the image features and the government-published poverty rates. The data derived from these features are In the final training step, ridge regression is implemented to determine the relationship between the aggregated by getting the element-wise average values of the vectors at the same geographic level as the government-published poverty rate. Ridge regression is linear like ordinary least squares regression, but it applies a squared penalty term (lambda) on the parameters to avoid overfitting in the case of a small ratio of observations to covariates. In principle, however, one may also consider using other model estimation methods like random forest to assess the sensitivity of estimates in the chosen estimation method.

STEP 1

For ridge regression, upload a new notebook file in Google Colab (footnote 7). Click **File**.

Click **Upload Notebook**.

STEP 3

Click **Choose File**.

Use the Jupyter Notebook file **Ridge_regression.ipynb**. Locate the file and click **Open**.

STEP 4

Click **Connect**. This will initialize the Colab environment.

Mount Google Drive (footnote 6) to Google Colab.

STEP 6

Click on the link.

STEP 7

In the browser, sign in to Google account.

Click **Allow**.

Click **Copy** icon $\boxed{\color{red}\Box}$ to copy the code.

Return to the Colab browser tab. **Paste** the code in the text box. Then press **Enter**.

A status will show the path where Google Drive is mounted.

STEP 8

Ensure that any edits made on the libraries are reloaded automatically and any charts or images displayed are shown in this notebook.

Locate the path to the dataset containing the binned luminosity and poverty rates.

Click **Files** icon **the Superior**.

STEP 10

Click on **gdrive** from the list of folders and expand the file directory tree to find the CSV file location.

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STEP 11

Click the vertical ellipsis to show more file options.

Click **Copy path**.

STEP 13

Paste the link on the blank space after the variable *CENI_full_file* and enclose in apostrophes.

Import the CSV file containing the merged training test dataset from Google Drive.

Define the different parameters needed for the model.

STEP 15

Drop all rows with "NA" values.

STEP 16

Load the feature dataset, which is the output of the feature extraction notebook, in the virtual machine drive. *It is then loaded as a dataframe.*

Compare the filenames of the daytime satellite imagery that were processed during feature extraction with the filename list from the original CSV file containing binned luminosity and poverty rates.

```
all img = features raw["filename"]
Œ
    # all img = pd. DataFrame(all img)
    missing images ID = df full['filename']. isin(all img)
    missing csventry ID = all img.isin(df full["filename"])
    missing images = df full[-missing images ID]
    missing entries = all img[-missing csventry ID]
    print("images in the df full, but not in the features file: ")
    print(missing images)
    print("^{\prime\prime}print("")print("images in the features file, but not in the df_full: ")
    print(missing entries)
[> images in the df_full, but not in the features file:
    Empty DataFrame
    Columns: [Unnamed: 0, id, lon, lat, geocode, avg_rad, bin_GMM, filename, Cit
    Index: []images in the features file, but not in the df_full:
    17868
             CNN DIMG PHI 2015 ST 384 3840 017869.jpg
    17869
             CNN DIMG PHI 2015 ST 384 3840 017870.jpg
             CNN DIMG PHI 2015 ST 384 3840 017941.jpg
    17940
             CNN_DIMG_PHI_2015_ST_384_3840_017942.jpg
    17941
    17942
             CNN DIMG PHI 2015 ST 384 3840 017943.jpg
    18016
             CNN_DIMG_PHI_2015_ST_384_3840_018017.jpg
    18786
             CNN DIMG PHI 2015 ST 384 3840 018787.jpg
             CNN_DIMG_PHI_2015_ST_384_3840_018788.jpg
    18787
             CNN DIMG PHI 2015 ST 384 3840 018789.jpg
    18788
             CNN_DIMG_PHI_2015_ST_384_3840_018836.jpg<br>CNN_DIMG_PHI_2015_ST_384_3840_018837.jpg
    18835
    18836
    18837
             CNN_DIMG_PHI_2015_ST_384_3840_018838.jpg
    18889
             CNN_DIMG_PHI_2015_ST_384_3840_018890.jpg
    18890
             CNN_DIMG_PHI_2015_ST_384_3840_018891.jpg
    18891
             CNN_DIMG_PHI_2015_ST_384_3840_018892.jpg
    18892
             CNN_DIMG_PHI_2015_ST_384_3840_018893.jpg
    18948
             CNN DIMG PHI 2015 ST 384 3840 018949.jpg
    18949
             CNN DIMG PHI 2015 ST 384 3840 018950.jpg
    18950
             CNN DIMG PHI 2015 ST 384 3840 018951.jpg
             CNN DIMG PHI 2015 ST 384 3840 019005.jpg
    19004
             CNN DIMG PHI 2015 ST 384 3840 019006.jpg
    19005
    19006
             CNN_DIMG_PHI_2015_ST_384_3840_019007.jpg
    Name: filename, dtype: object
```
Delete all rows in the original CSV file that contain filenames that were not processed during feature extraction.

STEP 18

Generate a new dataframe containing only the geocode and filenames column and drop duplicate geocode entries.

STEP 19

Generate a new dataframe containing only the training poverty data.

Merge the geocode-filename dataframe with the features dataframe.

```
# ensure that the datatypes align
img_geocode.filename.astype(str)
features_raw.filename.astype(str)
# merge
features = img geocode.merge(features raw, on = "filename")
```
STEP 21

Compute the average features by geocode group and generate one feature vector per geocode.

```
avg features = features.copy(deep = True)
avg features.drop(columns=['filename'])
avg_features = avg_features.groupby('geocode', as_index=False).mean()
```
STEP 22

Merge the training poverty and averaged features dataframes.

```
avg features full = df LHS.merge(avg features, on = 'geocode')
\blacktrianglerightprint(df LHS.shape)
   print(avg features.shape)
   print(avg_features_full.shape)
   print(avg_features_full.iloc[:5,:6])
\Gamma (1621, 3)
   (1621, 513)(1621, 515)geocode data_split POV_2015
                                                                    \overline{\mathbf{c}}0
                                                         1
   0 20902000
                                26.38 0.002491 0.001594 0.001437
                        0.91 20901000
                        0.914.40 0.001808
                                                  0.001810 0.001680
   2 20904000
                        0.117.96 0.001693 0.001549 0.001572
   3 20903000
                        0.918.27 0.002649 0.001554
                                                            0.003831
   4 20906000
                        0.919.48 0.001762 0.002304 0.001636
```
Load the packages needed to perform ridge regression.

STEP 24

Implement the following steps:

Determine geocodes of outliers from the averaged features based on the defined standard deviation specified in the variable *outlier_flag*.

```
Import numpy as np
    outlier flag = 4 \# standard deviation
    validation size percent = 10outliers = avg_features_full['geocode'][avg_features_full[target_variable_name] > avg_features_full[target_variable_name].mean() +
                                              outlier_flag * avg_features_full[target_variable_name].std()].unique()
    print("outlier Regions: ")
    print(outliers)
    print("number of outliers: " + str(len(outliers)))
   validation regions = avg features full['geocode'][avg features full['data split'] == (validation size percent/100)].unique()
    print("number of validation_regions: " + str(len(validation_regions)))
    # combine validation and outlier regions to drop them at once
    drop_regions = np.append(outliers, validation_regions)
    # drop outliers and validation set
    avg_features = avg_features_full[-avg_features_full['geocode'].isin(drop_regions)]
    avg_features_validation = avg_features_full[avg_features_full['geocode'].isin(validation_regions)]
   # training set
    Xs = avg_features.drop([target_variable_name, 'geocode', 'data_split'], axis = 1)
   y = avg_features[target_variable_name].values.reshape(-1,1)
    # full dataset
    Xs_full = avg_features_full.drop([target_variable_name, 'geocode', 'data_split'], axis = 1)
   y_full = avg_features_full[target_variable_name].values.reshape(-1,1)
    # only validation set
    Xs_validation = avg_features_validation.drop([target_variable_name, 'geocode', 'data_split'], axis = 1)
   y_validation = avg_features_validation[target_variable_name].values.reshape(-1,1)
   print(avg_features_full.shape)
   print("Xs shape: " + str(Xs.shape))<br>print("y shape: " + str(y.shape))
    print("Outlier flag: " + str(outlier_flag) + " sd")
   print("Validation Xs shape: " + str(Xs validation.shape))<br>print("Validation relative size: " + str(round( Xs_validation.shape[0] / avg_features_full.shape[0],2)) )
```
Extract the validation datasets and drop the outliers.

■ Create separate dataframes for full, training, and test datasets.

E outlier Regions: \mathbf{I} number of outliers: 0 number of validation regions: 161 $(1621, 515)$ Xs shape: (1460, 512) y shape: (1460, 1) Outlier flag: 4 sd Validation Xs shape: (161, 512) Validation relative size: 0.1

Set the parameter space for lambda (the ridge regression penalty term) that needs to be searched through.

```
\bullet max lambda = 10
    print("maximum lambda: " + str(max lambda))min lambda = 0.01print ("minimum lambda: " +str (min lambda))
    parameters = \{\text{ 'alpha': } 10**np.\text{linspace}(np.\text{log10}(min\_lambda), np.\text{log10}(max\_lambda), num = 15)\}print(parameters)
[-3 maximum lambda: 10
    minimum lambda: 0.01
    \{'alpha\}: array([ 0.01, 0.01637894, 0.02682696, 0.04393971, 0.07196857,
             0.11787686, \quad 0.19306977, \quad 0.31622777, \quad 0.51794747, \quad 0.8483429 \ \ ,1.38949549, 2.27584593, 3.72759372, 6.1054023, 10.
                                                                                  1)
```
STEP 26

Perform ridge regression.

```
D ridge = Ridge(fit_intercept = True, normalize = True)
    ridge_regressor = GridSearchCV(ridge, parameters, scoring = "neg_mean_squared_error", cv = 10)
    %time ridge_regressor.fit(Xs,y)
5 CPU times: user 9.21 s, sys: 9.27 s, total: 18.5 s
    Wall time: 9.4 s
    GridSearchCV(cv=10, error_score=nan,
                 estimator=Ridge(alpha=1.0, copy_X=True, fit_intercept=True,
                                  max_iter=None, normalize=True, random_state=None,
                                  solver='auto', tol=0.001),
                 iid='deprecated', n_jobs=None,
                 param_grid={'alpha': array([ 0.01
                                                           , 0.01637894, 0.02682696, 0.04393971, 0.07196857,
           0.11787686, 0.19306977, 0.31622777, 0.51794747, 0.8483429,<br>1.38949549, 2.27584593, 3.72759372, 6.1054023, 10. ))}
                 pre_dispatch='2*n_jobs', refit=True, return_train_score=False,
                  scoring='neg_mean_squared_error', verbose=0)
```
STEP 27

Identify the model with the best CV score.

```
print(ridge_regressor.best_params_)
   best_ridge = ridge_regressor.best_estimator_
   RMSE valid = round((y validation/100 - 0.01*)best ridge.predict(Xs validation))**2).mean(*)*0.5,4)RMSE full = round((y full/100 - 0.01*best ridge.predict(Xs full))*2).mean()*0.5,4)print("Validation RMSE: " + str(RMSE_valid))
   print("Full RMSE: " + str(RMSE_full))
   {'alpha': 0.8483428982440717}
   Validation RMSE: 0.1107
   Full RMSE: 0.1045
```
Define the function for computing R-squared and root mean square error (RMSE).

```
O import shutil
    def Ridge_Rsquared (predicted, true):
     SSE = sum((predicted - true)**2)SST = sum((true - true.mean())**2)
     R_square = 1 - SSE / SST
    RMSE = (SSE/len(true))**0.5
    return round(float(R_square),4)
    eval_valid = Ridge_Rsquared(0.01*best_ridge.predict(Xs_validation), 0.01*y_validation)
    eval_full = Ridge_Rsquared(0.01*best\_ridge.predict(Xs_full), 0.01*y_full))
    eval_train = Ridge_Rsquared(0.01*best\_ridge.predict(Xs), 0.01*y)ridgestats = pd.DataFrame({"stat": ['RMSE_valid', 'RMSE_full', 'R2_valid', 'R2_full', 'R2_train'],
                       "value": [RMSE_valid, RMSE_full, eval_valid, eval_full, eval_train]})
    print(ridgestats)
    ridgestats_file = "_".join(["CNN", "Ridgestats", "RES34", country, year, day_sat, str(img_res)] ) + ".csv"
    ridgestats.to_csv(ridgestats_file)
    shutil.copy(os.path.join('/content/',ridgestats_file), "/content/gdrive/MyDrive/")
            stat value
D
    0 RMSE_valid 0.1107
   1 RMSE full 0.1045
   2 R2 valid 0.5038
        R2 full 0.5972
    \overline{3}R2_train 0.6060
    4
    '/content/gdrive/MyDrive/CNN_Ridgestats_RES34_PHI_2015_ST_384.csv'
```
Implement the calculations for the training, validation, and the entire dataset.

```
shutil
    def Ridge_Rsquared (predicted, true):
     SSE = sum(</math> (predicted - true) **2)SST = sum((true - true mean)))**2)
      R square = 1 - SSE / SSTRMSE = (SSE/len(true))**0.5return round(float(R square), 4)
    eval_valid = Ridge_Rsquared(0.01*best_ridge.predict(Xs_validation), 0.01*y_validation)
    eval_full = Ridge_Rsquared(0.01*best\_ridge.predict(Xs_full), 0.01*y_full)eval train = Ridge Rsquared(0.01*best ridge.predict(Xs), 0.01*y)
    ridgestats = pd.DataFrame({"stat": ['RMSE_valid', 'RMSE_full', 'R2_valid', 'R2_full', 'R2_train'],
                       "value": [RMSE_valid, RMSE_full, eval_valid, eval_full, eval_train]})
    print(ridgestats)
    ridgestats_file = "_".join(["CNN", "Ridgestats" , "RES34", country, year, day_sat, str(img_res)] ) + ".csv"
    ridgestats.to_csv(ridgestats_file)
    shutil.copy(os.path.join('/content/',ridgestats_file), "/content/gdrive/MyDrive/")
            stat value
C.
    0 RMSE_valid 0.1107
    \mathbf 1RMSE_full 0.1045
   2 R2 valid 0.5038
    \overline{\mathbf{3}}R2_full 0.5972
       R2 train 0.6060
    4
    '/content/gdrive/MyDrive/CNN_Ridgestats_RES34_PHI_2015_ST_384.csv'
```
Generate the regression statistics outputs as CSV file and copy them in Google Drive.

```
O import shutil
    def Ridge_Rsquared (predicted, true):
     SSE = sum((predicted - true)**2)
     SST = sum((true - truemean())**2)
      R square = 1 - SSE / SSTRMSE = (SSE/len(true))**0.5
     return round(float(R_square), 4)
    eval_valid = Ridge_Rsquared(0.01*best_ridge.predict(Xs_validation), 0.01*y_validation)
    eval_full = Ridge_Rsquared(0.01*best_ridge.predict(Xs_full), 0.01*y_full)
    eval_train = Ridge_Rsquared(0.01*best_ridge.predict(Xs), 0.01*y)
    ridgestats = pd.DataFrame({"stat": ['RMSE_valid', 'RMSE_full', 'R2_valid', 'R2_full', 'R2_train'],
                       "value": [RMSE_valid, RMSE_full, eval_valid, eval_full, eval_train]})
    print(ridgestats)
    ridgestats_file = "_".join(["CNN", "Ridgestats", "RES34", country, year, day_sat, str(img_res)] ) + ".csv"
    ridgestats.to_csv(ridgestats_file)
    shutil.copy(os.path.join('/content/',ridgestats_file), "/content/gdrive/MyDrive/")
            stat value
Ŀ
   0 RMSE_valid 0.1107
   1 RMSE_full 0.1045
   2 R2_valid 0.5038
        R2 full 0.5972
   \overline{3}R2_train 0.6060
    '/content/gdrive/MyDrive/CNN_Ridgestats_RES34_PHI_2015_ST_384.csv'
```
STEP 31

Import the **matplotlib** library used for data visualization. Then define a function for plotting a 45-degree fit line.

```
import matplotlib.pyplot as plt
\mathbf{D}# add functionality to plot at 45° line
    def abline(slope, intercept):
        """Plot a line from slope and intercept"""
        axes = plt.get()x vals = np.array(axes.get_xlim())y vals = intercept + slope * x vals
        plt.plot(x_values, y_values, '-')
```
Plot the government-published poverty rates against the predicted poverty rates.

STEP 33

Load the Python **pickle** library, which then exports the ridge regression model. Copy the file to Google Drive.

STEP 34

Then reload the saved model parameters.

Extract the array of the image level features, collapse it into a one-dimension array to get the predicted poverty rates, and generate a dataframe with the corresponding imagery filename as the index.

Then, merge the poverty prediction dataframe with the data frame containing the government-published poverty rates using the imagery filename as the merging parameter.

STEP 36

Generate the poverty prediction output file as a CSV file. Then copy these results to Google Drive.

```
Q
   poverty prediction file = " ".join(["CNN", "POV" , "RES34", country, year, day sat, str(img res)] ) + ".csv"
   output.to_csv(poverty_prediction_file)
    shutil.copy(os.path.join('/content/',poverty_prediction_file), "/content/gdrive/MyDrive/")
    "/content/gdrive/MyDrive/CNN_POV_RES34_PHI_2015_ST_384.csv"
```
RESCALING OF **7 RESCALING OF POVERTY ESTIMATES** AND VISUALIZATION

Data Requirements

- CSV file containing poverty estimates
- **Machine learning based population estimate raster**

Tool

■ R and RStudio (Rescaling_and_visualization.R)

STEP 1

In RStudio, use the R code: **Rescaling_and_visualization.R.**

continued on next page

Step 1 *continued*

STEP 2

Load **raster** and **tidyverse** packages.

STEP 3

Define the folder where the temporary raster files will be saved or create the folder if it does not exist.

For raster calculations, set several raster package options to improve the speed of calculation. The important options are as follows:

- **number 1 maxmemory** maximum number of bytes to read into memory.
- **chunksize** maximum number of bytes to read/write in a single chunk while processing (chunk by chunk) disk-based raster objects.

Other options are:

- **progress** 'text': displays raster operation progress bar
- **tmptime** number of hours before a temporatry file gets deleted from the tmpdir.
- **tmpdir** location for writing temporary file.
- **timer** TRUE: outputs the raster calculation duration.

```
tmp_path <- "C:/temp"
 \overline{7}8
9 - if (ldir.exists(tmp.path))dir.create(tmp_path)
                                    # Create the folder if not yet existing
10
11 + 31213 - #Set raster options ----
14 raster0ptions(tmptime = 4,
15
                  progress = 'text',16
                  timer = TRUE,17
                  maxmemory = 10e+9,
18
                  chunksize = 5e+9,
19
                  tmpdir = tmp_path)
```
Define the coordinate reference system for WGS84.

```
\overline{c}\overline{v}21 - # Define CRS----
22 WGS84<- "+proj=longlat +datum=WGS84 +no_defs +ellps=WGS84 +towgs84=0,0,0"
23
```
STEP 5

Select the CSV file containing the ridge regression poverty estimates.

```
24 # select csv file containing the ridge regression predicted poverty<br>25 pov_csv_path <- tcltk::tk_choose.files(filters = matrix(c("CSV",".csv","All files","*"),2,2,byrow = T),<br>26 caption = "Select Predicted Poverty CSV"
```


Set the CSV's parent directory as the working directory. Extract the country code and year of study using information from the CSV filename. Then, define the government-published poverty estimates' column name. Load the CSV file as a dataframe.

STEP 7

Subset the predicted poverty dataframe to get the grid ID (id) and the latitude (lat) and longitude (lon), and rasterize the resulting dataframe using the function *rasterFromXYZ( )*.

The function *rasterFromXYZ( )* generates raster from regular grids like the dataset used. The function assumes that the minimum distance between x and y coordinates is the raster resolution.

STEP 8

Load the machine learning population raster.

```
51 + #load ML estimated population raster----
52 pop_raster_path <- tcltk::tk_choose.files(filters = matrix(c("TIF",".tif","All files","*"),2,2,byrow = T)
                                         caption = "Select Population Raster")
53
54 pop_raster <- raster(pop_raster_path)
```


Check if the population raster is using WGS84 CRS. Otherwise, reproject the raster. Print out the new CRS of the population raster. Also, compare the resolution of the population and poverty grids. Note from the results that the population and centroid rasters have different resolutions.

```
56 #check if pop_raster projection is NGS84, otherwise reproject raster
57 - if (compareCRS(pop_raster,WGS84) == FALSE) {
58 print("Raster CRS is not WGS84. Projecting raster to WGS84...")
59
    pop_raster <- projectRaster(pop_raster,crs = WGS84)
60 - 761 print(crs(pop_raster))
62 print(paste0("Population Raster grid size: ",paste(res(pop_raster),collapse = ", ")))
63 print(paste0("Centroid Raster grid size: ",paste(res(centroid_rast),collapse = ", ")))
```

```
> #check if pop_raster projection is WGS84, otherwise reproject raster
> if (compareCRS(pop_raster,WGS84)==FALSE) {
  print("Raster CRS is not WGS84. Projecting raster to WGS84...")
+pop_raster <- projectRaster(pop_raster,crs = WGS84)
++3[1] "Raster CRS is not WGS84. Projecting raster to WGS84..."
361 seconds
> print(crs(pop_raster))
CRS arguments:
+proj=longlat +datum=WGS84 +no_defs +ellps=WGS84 +towgs84=0,0,0
> print(paste0("Population Raster grid size: ",paste(res(pop_raster),collapse = ", ")))
[1] "Population Raster grid size: 0.000921, 0.000903000000000003"
> print(paste0("Centroid Raster grid size: ",paste(res(centroid_rast),collapse = ", ")))
[1] "Centroid Raster grid size: 0.0429204653119797, 0.0429204653119974"
\geq
```
Calculate the **adjustment_factor** first because the two rasters have different resolutions.

Aggregate the population headcount of the machine learning population raster at the poverty grid level, which will be used to rescale the ridge regression poverty prediction. Using the **aggregate( )** function, aggregate the population in the poverty grid using the calculated adjustment_factor. Then, resample the aggregated population raster to match the resolution of the centroid raster.

```
65 # determine resolution ratio of centroid raster and population raster
67
68 # aggregate population raster values to poverty grid by taking its sum
69 pop_agg <- aggregate(pop_raster, fact = adjustment_factor, fun = sum)
70
71 # resample pop_agg raster to match the extent and resolution of centroid_rast
72 pop_agg_resampled <- resample(pop_agg, centroid_rast)
```
STEP 11

Set the aggregated population raster layer's name to *"gridpop"*. Stack the centroid and aggregated population raster, then convert the raster stack as a dataframe. Merge the created dataframe with the predicted poverty dataframe.

```
# rename raster column
74
75 names(pop_agg_resampled) <- "gridpop"
76
77 # stack the two raster
78 pop_id_stack <- raster::stack(centroid_rast,pop_agg_resampled)
79
80 #convert the raster stack to dataframe
81 df_pop_id <- as.data.frame(pop_id_stack,na.rm=T)
82
83 # merge the aggregated population at poverty grid with the predicted poverty dataframe
84 df_grid_pov <- left_join(df_pov,df_pop_id,by="id")
```
STEP 12

Prior to rescaling, check if there are poverty prediction values that are either negative or more than 100%. Set the negative values to 0.0001 and adjust the values above 100% to 100%.

```
86 + # Rescaling poverty estimates ----
87
88 #list predictions with values less than zero
89 df_grid_pov$prediction[df_grid_pov$prediction<0]
90 #list predictions with values more than 100
91 df_grid_pov$prediction[df_grid_pov$prediction>100]
92
93 # set all negative and more than 100 prediction values to 0.0001 and 100, respectively
94 df_grid_pov$prediction[df_grid_pov$prediction<0] <- 0.0001
95 df_grid_pov$prediction[df_grid_pov$prediction>100] <- 100
```


Rescale the poverty predictions. Convert the predicted poverty rates to index values by dividing the values by 100.

```
97
     #rescale poverty predictions based on published poverty estimates
98
    df grid pov \lt- df grid pov \frac{8}{5}99
       mutate(pred_hci = prediction / 100) %100
       mutate(svy_hci = get(target_var) / 100) %101
       mutate(pred_hc = gridpop * pred_hci) %>\%102
       mutate(svy_hc = gridpop * svy_hci) %103
       group_by(geocode) %>%
104
       mutate(pred_hc_rescale = pred_hc * (sum(svy_hc) / sum(pred_hc))) %>%
105
       mutate(pred_hci_rescale = pred_hc_rescale / gridpop) %>%
106
       ungroup()
```
Convert the government-published poverty rates to index values by dividing the values by 100.

STEP 15

Calculate the grid level poverty headcount by multiplying the grid population by the predicted poverty index.

```
97
     #rescale poverty predictions based on published poverty estimates
 98
     df_grid_pov <- df_grid_pov %>%
 99
       mutate(pred_hci = prediction / 100) %100
       mutate(svy_hci = get(target_var) / 100) %101
       mutate(pred_hc = gridpop * pred_hci) %>\102
       mutate(svy_hc = gridpop * svy_hci) %>%
103
       group_by(geocode) %>%
104
       mutate(pred_hc_rescale = pred_hc * (sum(svy_hc) / sum(pred_hc))) %>%
105
       mutate(pred_hci_rescale = pred_hc_rescale / gridpop) %>%
106
       ungroup()
```
STEP 16

Calculate the government-published poverty headcount.

```
97
     #rescale poverty predictions based on published poverty estimates
     df_grid_pov <- df_grid_pov %>%
 98
 99
       mutate(pred_hci = prediction / 100) %100
       mutate(svy_hci = get(target_var) / 100) %101
       mutate(pred_hc = gridpop * pred_hci) %>\%102
       mutate(svy_hc = gridpop * svy_hci) %103
       group_by(geocode) %>%
104
       mutate(pred_hc_rescale = pred_hc * (sum(svy_hc) / sum(pred_hc))) %>%
105
       mutate(pred_hci_rescale = pred_hc_rescale / gridpop) %>%
106
       ungroup()
```
Group the data according to geocode.

STEP 18

Derive the rescaled predicted poverty headcount for each grid by multiplying the grid's predicted poverty headcount by the ratio of the sum of the government-published and predicted poverty headcounts. *This is calculated for each geocode group.*

```
97
     #rescale poverty predictions based on published poverty estimates
     df_{\text{grid}pov} \leftarrow df_{\text{grid}pov} %>%
 98
99
       mutate(pred_hci = prediction / 100) %>100
       mutate(svy_hci = get(target_var) / 100) %101
       mutate(pred_hc = gridpop * pred_hci) %>\102
       mutate(svy_hc = gridpop * svy_hci) %103
       group_by(geocode) %>%
104
       mutate(pred_hc_rescale = pred_hc * (sum(svy_hc) / sum(pred_hc))) %>%
105
       mutate(pred_hci_rescale = pred_hc_rescale / gridpop) %>\n106
       ungroup()
```
STEP 19

Calculate the rescaled poverty index by dividing the rescaled predicted poverty headcount by the grid level population counts.

```
97
     #rescale poverty predictions based on published poverty estimates
 98
     df_grid_pov <- df_grid_pov %>%
       mutate(pred_hci = prediction / 100) %99
100
       mutate(svy_hci = get(target_var) / 100) %101
       mutate(pred_hc = gridpop * pred_hci) %>\%102
       mutate(svy_hc = gridpop * svy_hci) %>>103
       group_by(geocode) %>%
       mutate(pred_hc_rescale = pred_hc * (sum(svy_hc) / sum(pred_hc))) %>%
104
105
       mutate(pred_hci_rescale = pred_hc_rescale / gridpop) %>%
106
       ungroup()
```
Ungroup the dataframe.

```
97
     #rescale poverty predictions based on published poverty estimates
 98 df_grid_pov <- df_grid_pov %>%
99
       mutate(pred_hci = prediction / 100) %100
       mutate(svy_hci = get(target_var) / 100) %101
       mutate(pred_hc = gridpop * pred_hci) %>\%102
       mutate(svy_hc = gridpop * svy_hci) %103
       group_by(geocode) %>%
104
       mutate(pred_hc_rescale = pred_hc * (sum(svy_hc) / sum(pred_hc))) %>%
       mutate(pred_hci_rescale = pred_hc_rescale / gridpop) %>%
105
106
       ungroup()
```
STEP 21

Check if there are rescaled poverty indexes above 1; set to 1 if there are any.

```
#list rescaled predictions with values more than 1
101
102
     df_grid_pov$pred_hci_rescale[df_grid_pov$pred_hci_rescale>1]
103
104
     # If any, set all rescaled values more than 1 to 1
     df_grid_pov$pred_hci_rescale[df_grid_pov$pred_hci_rescale>1] <- 1
105
```


Generate the poverty raster.

```
115 - # generate raster
116
117 pov_hci_raster <- rasterFromXYZ(xyz = df_grid_pov[,c("lon","lat", "pred_hci")], crs=WGS84)
118 pov_hci_rescaled_raster <- rasterFromXYZ(xyz = df_grid_pov[,c("lon","lat", "pred_hci_rescale")], crs=WGS84)
119
120 - # Output raster----
121 # set raster destination path
122 raster_path <- "Output/Poverty Raster/"
123
124 - if (!dir.exists(raster_path)) {
125
     dir.create(raster.path, recursive = T)126 - 1127
128 writeRaster(pov_hci_raster,
129
                 filename = paste0(raster_path, paste(country_year, "pov_hci.tif",sep = "_")),
130
                 overwrite=TRUE)
131 writeRaster(pov_hci_rescaled_raster,
132
                 filename = paste0(raster_path, paste(country_year,"pov_hci_rescaled.tif",sep = "_")),
133
                 overwrite=TRUE)
```
Generate poverty rasters for both predicted and rescaled predicted poverty index using the raster function **rasterfromXYZ( )**.

The parameters supplied are the centroid coordinates (lat and lon) and the corresponding data to be rasterized.

STEP 23

Define the folder where the raster will be saved or create the folder if it does not exist.

```
115 - # generate raster --
116
     pov_hci_raster <- rasterFromXYZ(xyz = df_grid_pov[,c("lon","lat", "pred_hci")], crs=WGS84)
117
118
     pov_hci_rescaled_raster <- rasterFromXYZ(xyz = df_grid_pov[,c("lon","lat", "pred_hci_rescale")], crs=WGS84)
119
120 - # Output raster---
121 # set raster destination path
122
     raster_path <- "Output/Poverty Raster/"
123
124 -if (!dir.exists(raster_path)) {
125
      dir.create(raster.path, recursive = T)126 -127
128
     writeRaster(pov_hci_raster,
129
                 filename = paste0(raster_path, paste(country_year, "pov_hci.tif",sep = "_")),
130
                 overwrite=TRUE)
131 writeRaster(pov_hci_rescaled_raster,
                 filename = paste0(raster_path, paste(country_year,"pov_hci_rescaled.tif",sep = "_")),
132
133
                 overwrite=TRUE)
```
Output the raster using the *writeRaster( )* function.

```
115 - # generate raster -
116
117 pov_hci_raster <- rasterFromXYZ(xyz = df_grid_pov[,c("lon","lat", "pred_hci")], crs=WGS84)
118 pov_hci_rescaled_raster <- rasterFromXYZ(xyz = df_grid_pov[,c("lon","lat", "pred_hci_rescale")], crs=WGS84)
119
120 - # Output raster----
121 # set raster destination path
122 raster_path <- "Output/Poverty Raster/"
123
124 - if (!dir.exists(raster_path)) {
125
       dir.create(raster.path, recursive = T)126 -}
127
128
     writeRaster(pov_hci_raster
129
                 filename = paste0(raster_path, paste(country_year, "pov_hci.tif",sep = "_")),
130
                 overwrite=TRUE)
131
     writeRaster(pov_hci_rescaled_raster,
132
                 filename = paste0(raster_path, paste(country_year,"pov_hci_rescaled.tif",sep = "_")),
                 overwrite=TRUE)
133
```
STEP 25

Visualize the raster. Load another raster visualization package, **rasterVis** (aside from **ggplot2**, which was already loaded as part of the **tidyverse** package).

```
135 - # Visualization--
136
     #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 = plot_raster <- function(rast, p_var){
141
        theme_set(theme_bw())
        hci_heat <- cut(rast, p_var$category/100, include.lowest = T)
142
143
144
        plt\_raster \leftarrow qplot(hci\_heat) +145
          geom\_tile(aes (fill = as.charAt(cvalue))) +146
          scale_fill_brewer(name = p_var$scale_title,
                              palette = "RdYlGn",147
                              direction = -1,
148
149
                              labels = p_{var$scale\_label} +
150
          labs( title = paste0(p_var$map.title),x =<sup>""</sup>,
151
                 y = m + 1152
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_\mathsf{blank}(),154
                 axis. ticks = element_bland(),panel.get.major = element_blank(),155
156
                 panel.get.d.minor = element_blank(),
                 panel.border = element_blank())+
157
158
          coord_fixed()
```
Define **plot_raster()** function that will aid in plotting the raster.

```
135 - # Visualization---136 #load packages
137
     library(rasterVis)
138
139
      #define plotting function
140 -plot_raster <- function(rast,p_var){
141
        theme_set(theme_bw())
        hci_{\text{heat}} \leftarrow \text{cut}(\text{rast}, p_{\text{water}}\text{scotegory}/100, \text{include}.\text{lowest} = T)142
143
144
        plt\_raster \leftarrow gplot(hci\_heat) +145
          geom_tile(aes(fill = as.charAter(value))) +146
          scale_fill_brewer(name = p_var$scale_title,
147
                               palette = "RdYlGn",148
                               direction = -1,149
                               labels = p_{var$scale\_label} +
150
          labs( title = paste0(p_var$map_title),151
                 x = "",
                 y = m +
152
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_\mathsf{blank}(),154
                 axis. ticks = element_blank(),155
                 panel. grid.major = element_blank(),156
                 panel.getd.minor = element_blank(),
157
                 panel.border = element_blank()+
158
          coord_fixed()
159
160
        #save map as png
161
        ggsave(plt_raster,
162
                filename = p_{var$filtername},163
                dpi = 300,
164
                device='png')165
166
        return(plt_raster)
167 -
```
The function requires two objects, a raster (**rast**) and a list (**p_var**). *p_var* contains the following parameters:

- **a category** a vector object containing the interval classes for reclassifying the raster values,
- *scale_title* and *scale_label* define the scale bar title and labels, respectively,
- *map_title* defines the map title, and
- **Filename** specifies the filename of the map for saving as png image file.

Inside the function, set the theme to black and white.

```
135 - # Visualization---136
     #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 * plot_raster <- function(rast, p_var){
141
        theme_set(theme_bw())
142
        hci_{\text{} -}heat <- cut(rast, p_var$category/100, include.lowest = T)
143
144
        plt\_raster \leftarrow gplot(hci\_heat) +145
          geom\_tile(aes(fill = as.charAtacter(value))) +146
          scale_fill_brewer(name = p_var$scale_title,
147
                              palette = "RdYlGn",148
                              direction = -1,labels = p_{var$scale\_label} +
149
150
          labs( title = paste0(p_var$map.title),x = \frac{mn}{n}151
                y ="") +
152
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_\mathsf{blank}(),axis. ticks = element_blank(),154
                 panel.get.major = element_blank(),155
156
                 panel.get.d.minor = element_blank(),157
                 panel.border = element_blank() +
158
          coord_fixed()
159
160
        #save map as png
161
        ggsave(plt_raster,
               filename = p_{var$filtername},162
163
               dpi = 300,
               device='png')
164
165
166
        return(plt_raster)
167 -}
```
Using the supplied category, reclassify the raster values

```
135 - # Visualization---136 #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 * plot_raster <- function(rast, p_var){
141
       theme_set(theme_bw())
142
       hci_heat <- cut(rast, p_var$category/100, include.lowest =
                                                                    \top143
144
       plt\_raster \leftarrow gplot(hci\_heat) +geom_tile(aes(fill = as.character(value)))+
145
146
         scale_fill_brewer(name = p_var$scale_title,palette = "RdYlGn",147
148
                            direction = -1,
                            labels = p_{var$scale\_label} +
149
150
         labs( title = paste0(p_var$map.title),151
               x = \frac{m n}{2}y ="") +
152
153
         \text{theme}(axis.text = element\_blank(),154
               axis. ticks = element_blank(),panel.get.major = element_blank(),155
               panel.get.minor = element_blank(),156
                panel.border = element_blank() +
157
158
         coord_fixed()
159
160
       #save map as png
161
       ggsave(plt_raster,
162
               filename = p_{var$filtername},163
              dpi = 300,
164
              device='png')
165
166
       return(plt_raster)
167 -}
```
Create a **gplot** object and set the categorized raster as the data source. *gplot* is a wrapper for plotting raster.

```
135 - # Visualization---136
     #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 - plot_raster <- function(rast, p_var){
141
        theme_set(theme_bw())
142
        hci_{\text{} +} heat <- cut(rast, p_var$category/100, include.lowest = T)
143
        plt\_raster \leftarrow glob(hci\_heat) +144
145
          geom_tile(aes(fill = as.character(value)))+
146
          scale_fill_brewer(name = p_var$scale_title,
147
                              palette = "RdYlGn",148
                              direction = -1,labels = p_{var$scale\_label} +
149
150
          labs( title = paste0(p_var$map.title),x = \frac{m n}{n}151
                y = m +
152
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_\mathsf{blank}(),axis. ticks = element_blank(),154
                panel.get.d.major = element_blank(),155
156
                 panel.get.d.minor = element_blank(),157
                 panel.border = element_bland()+
158
          coord_fixed()
159
160
        #save map as png
161
        ggsave(plt_raster,
162
               filename = p_{var$filtername},dpi = 300,
163
164
               device='png')165
166
        return(plt_raster)
167 -}
```
Specify the raster's value as the object fill using the **geom_tile( )** function.

```
135 - # Visualization---136
     #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 - plot_raster <- function(rast, p_var){
141
        theme_set(theme_bw())
142
        hci_heat <- cut(rast, p_var$category/100, include.lowest = T)
143
144
        plt\_raster \leftarrow gplot(hci\_heat) +145
          geom\_tile(aes(fill = as.charAtacter(value))) +146
          scale_fill_brewer(name = p_var$scale_title,
                             palette = "RdYlGn",
147
148
                             direction = -1,149
                             labels = p_{var$scale\_label} +
150
          labs( title = paste0(p_var5map_title),151
                x =<sup>""</sup>,
                y = m + 1152
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_ \mathsf{blank}(),154
                axis.ticks = element_blank(),
                panel.grid.major = element_blank(),
155
156
                panel.get.d.minor = element_blank(),157
                panel.border = element_blank()+
          coord_fixed()
158
159
160
        #save map as png
161
        ggsave(plt_raster,
162
               filename = p_var$filename,163
               dpi = 300,
164
               device='png')165
166
        return(plt_raster)
167 -}
```
Using the *scale_fill_brewer( )* function, specify the following:

- **name** scale title,
- **palette** color palette of the map and scale, which is set to Red-Yellow-Green ("RdYIGn"),
- **direction = -1** reverses the color palette order from "RdYlGn" to "GnYlRd", and
- **labels** scale label to match the categorical grouping of the dataset.

```
135 - # Visualization---136
     #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 - plot_raster <- function(rast, p_var){
141
        theme_set(theme_bw())
142
        hci_heat <- cut(rast, p_var$category/100, include.lowest = T)
143
144
        plt_raster <- gplot(hci_heat) +
145
          geom\_tile(aes (fill = as.charAt (value))) +146
          scale_fill_brewer(name = p_var$scale_title,
147
                             palette = "RdYlGn",148
                             direction = -1.
149
                             labels = p_{var$scale\_label} +
          labs(title = paste0(p_var$map_title),150
151
                x = \frac{m n}{2}152
                y = '''') +
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_\mathsf{blank}(),154
                axis. ticks = element_blank(),155
                panel.get.d.major = element_blank(),156
                panel. grid.minor = element_blank().
157
                panel.border = element_blank() +
          coord_fixed()
158
159
160
        #save map as png
161
        ggsave(plt_raster,
162
               filename = p_{var$filtername},dpi = 300,
163
164
               device='png')165
166
        return(plt_raster)
167 -}
```
Specify the map title and leave the x and y axes unlabeled.

```
135 - # Visualization---136 #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 - plot_raster <- function(rast, p_var){
141
       theme_set(theme_bw())
142
       hci\_heat \leftarrow cut(rast, p\_var$category/100, include.lower = T)143
144
       plt\_raster \leftarrow glob(hci\_heat) +145
         geom\_tile(aes (fill = as.charAter(value))) +146
         scale_fill_brewer(name = p_var$scale_title,147
                            palette = "RdYlGn",148
                            direction = -1,149
                            labels = p_{var$scale\_label} +
150
         labs( title = paste0(p_varsmap_title),x = \frac{m}{n}151
               y = m +
152
         theme(axis.text = element_blank(),153
154
               axis. ticks = element_blank(),155
               panel.get.d.major = element_blank(),156
               panel.get.d.minor = element_blank(),panel.border = element_blank() +
157
158
         coord_fixed()
159
160
       #save map as png
161
       ggsave(plt_raster,
162
              filename = p_var$filename,163
              dpi = 300,
164
              device='png')
165
166
       return(plt_raster)
167 -}
```
Remove axis text, tick marks, gridlines, and borders (optional).

```
135 - # Visualization---136
     #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 - plot_raster <- function(rast, p_var){
141
        theme_set(theme_bw())
        hci_{\text{} -}heat <- cut(rast, p_var$category/100, include.lowest = T)
142
143
        plt\_raster \leftarrow qplot(hci\_heat) +144
145
          geom\_tile(aes(fill = as.charAtarcter(value))) +146
          scale_fill_brewer(name = p_var$scale_title,
147
                              palette = "RdYlGn",148
                              direction = -1,
149
                              labels = p_{var$scale\_label} +
150
          labs( title = paste0(p_var\$map_title),x =<sup>""</sup>,
151
                y = m +152
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_ \mathsf{blank}(),154
                axis.ticks = element_blank(),
                panel.get.major = element_blank(),155
156
                 panel.get.d.minor = element_blank(),
157
                panel.border = element_blank()+
158
          coord_fixed()
159
160
        #save map as png
161
        ggsave(plt_raster,
162
               filename = p_{var$filtername},163
               dpi = 300,
164
               device='png')165
166
        return(plt_raster)
167 -}
```
Set the Cartesian coordinates to a fixed aspect ratio (**coord_fixed( )**) which is a 1:1 ratio of x and y values.

```
135 - # Visualization---136 #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 * plot_raster <- function(rast, p_var){
141
        theme_set(theme_bw())
142
        hci_heat <- cut(rast, p_var$category/100, include.lowest = T)
143
144
        plt\_raster \leftarrow gplot(hci\_heat) +145
          geom\_tile(aes(fill = as.charAtacter(value))) +146
          scale_fill_brewer(name = p_var$scale_title,
147
                             palette = "RdYlGn",148
                             direction = -1,149
                             labels = p_{var$scale\_label} +
150
          labs( title = paste0(p_var$map_title),151
                x = \frac{m n}{2}y = m + 1152
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_\mathsf{blank}(),154
                axis. ticks = element_blank(),155
                panel.get.major = element_blank(),156
                panel.get.d.minor = element_blank(),
                panel.border = element_blank()+
157
158
          coord_fixed()
159
160
        #save map as png
161
        ggsave(plt_raster,
162
               filename = p_{var$filtername},163
               dpi = 300,
164
               device='png')165
166
        return(plt_raster)
167 -}
```
Save the plot as png image format using the filename to be supplied in the variable **p_var**.

Other supported image format are "eps", "ps", "tex" (pictex), "pdf", "jpeg", "tiff", "png", "bmp", "svg" or "wmf".

```
135 - # Visualization---136
     #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 * plot_raster <- function(rast, p_var){
141
        theme_set(theme_bw())
        hci_heat <- cut(rast, p_var$category/100, include.lowest = T)
142
143
        plt\_raster \leftarrow glob(hci\_heat) +144
145
          geom_tile(aes(fill = as.charAter(value))) +scale_fill_brewer(name = p_var$scale_title,
146
147
                             palette = "RdYlGn",148
                             direction = -1,149
                             labels = p_{var$scale\_label} +150
          labs( title = paste0(p_var$map.title),x = \frac{mn}{n}151
                y ="") +
152
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_\mathsf{blank}(),153
                axis. ticks = element_blank(),154
155
                panel.get.d.major = element_blank(),156
                panel.get.d.minor = element_blank(),panel.border = element_blank() +
157
          coord_fixed()
158
159
160
        #save map as png
161
        ggsave(plt_raster,
               filename = p_{var$filtername},162
163
               dpi = 300,
164
               device='png')165
166
        return(plt_raster)
167 -}
```
Return the gplot object so that it will automatically show in the viewer pane upon function call.

```
135 - # Visualization---136 #load packages
137
     library(rasterVis)
138
139
     #define plotting function
140 * plot_raster <- function(rast, p_var){
141
       theme_set(theme_bw())
142
       hci_heat <- cut(rast, p_var$category/100, include.lowest = T)
143
144
       plt_raster <- gplot(hci_heat) +
145
          geom_tile(aes(fill = as.charAter(value))) +146
          scale_fill_brewer(name = p_var$scale_title,
147
                             palette = "RdYlGn",148
                             direction = -1,149
                             labels = p_{var$scale\_label} +
150
          labs( title = paste0(p_var5map_title),x = \frac{mn}{n}151
                y = m + 1152
153
          \mathsf{theme}(\mathsf{axis}.\mathsf{text} = \mathsf{element}\_\mathsf{blank}(),154
                axis. ticks = element_blank(),
155
                panel.get.d.major = element_blank(),156
                panel.getd.minor = element_blank(),157
                panel.border = element_blank()+
158
          coord_fixed()
159
160
       #save map as png
161
        ggsave(plt_raster,
               filename = p_{var$filtername},162
               dpi = 300,
163
164
               device='png')
165
166
       return(plt_raster)
167 -}
```
Set the maps' save path and create a folder if it does not exist.

STEP 37

Specify the parameters needed by the function and pass on the raster object and the parameters to the function.

The resulting poverty maps—machine learning (predicted and calibrated) and government-published for the Philippines are shown in Figure 2 and for Thailand in Figure 3.

Figure 2: Machine Learning and Published Poverty Rate Maps of the Philippines, 2015

Note: The first two images present the uncalibrated and calibrated machine learning-based poverty rate estimates in (approximately) every 4 square kilometer grid, respectively. The third image shows the municipal or city-level poverty rates published by the Philippine Statistics Authority.

Source: Calculations and graphics generated by the study team.

Figure 3: Machine Learning and Published Poverty Rate Maps of Thailand, 2015

Note: The first two images present the uncalibrated and calibrated machine learning-based poverty rate estimates in (approximately) every 4 square kilometer grid, respectively. The third image shows the tambon-level poverty rates published by the National Statistical Office of Thailand.

Source: Calculations and graphics generated by the study team.

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A Guidebook on Mapping Poverty through Data Integration and Artificial Intelligence

The "leave no one behind" principle of the 2030 Agenda for Sustainable Development requires appropriate indicators to be estimated for different segments of a country's population. The Asian Development Bank, in collaboration with the Philippine Statistics Authority, the National Statistical Office of Thailand, and the World Data Lab, conducted a feasibility study that aimed to enhance the granularity, cost-effectiveness, and compilation of high-quality poverty statistics in the Philippines and Thailand. This accompanying guide to the *Key Indicators for Asia and the Pacific 2020* special supplement is based on the study, capitalizing on satellite imagery, geospatial data, and powerful machine-learning algorithms to augment conventional data collection and sample survey techniques.

About the Asian Development Bank

ADB is committed to achieving a prosperous, inclusive, resilient, and sustainable Asia and the Pacific, while sustaining its efforts to eradicate extreme poverty. Established in 1966, it is owned by 68 members —49 from the region. Its main instruments for helping its developing member countries are policy dialogue, loans, equity investments, guarantees, grants, and technical assistance.

ASIAN DEVELOPMENT BANK

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